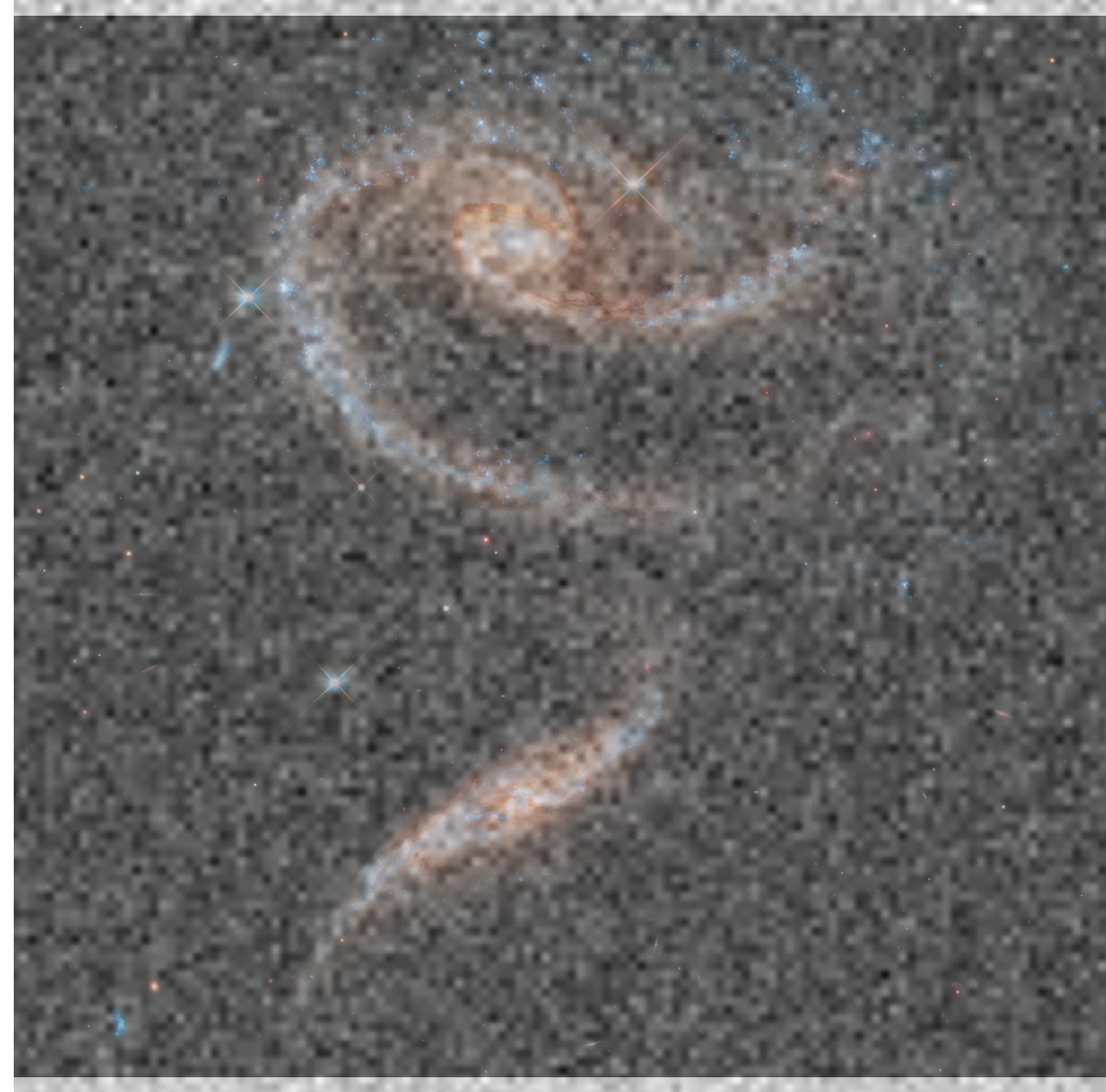


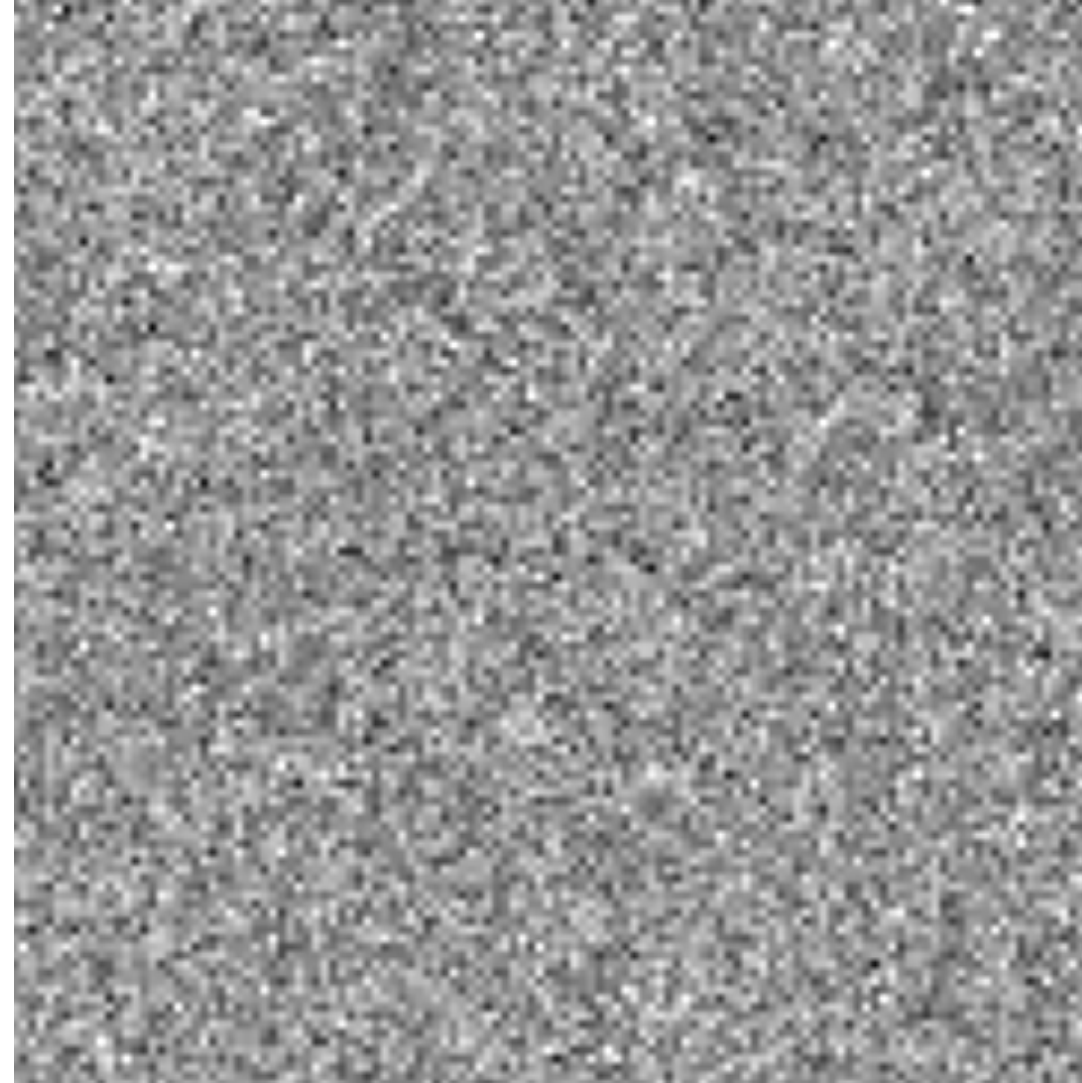
# Overcoming inference challenges using score generative models

**Ronan Legin**



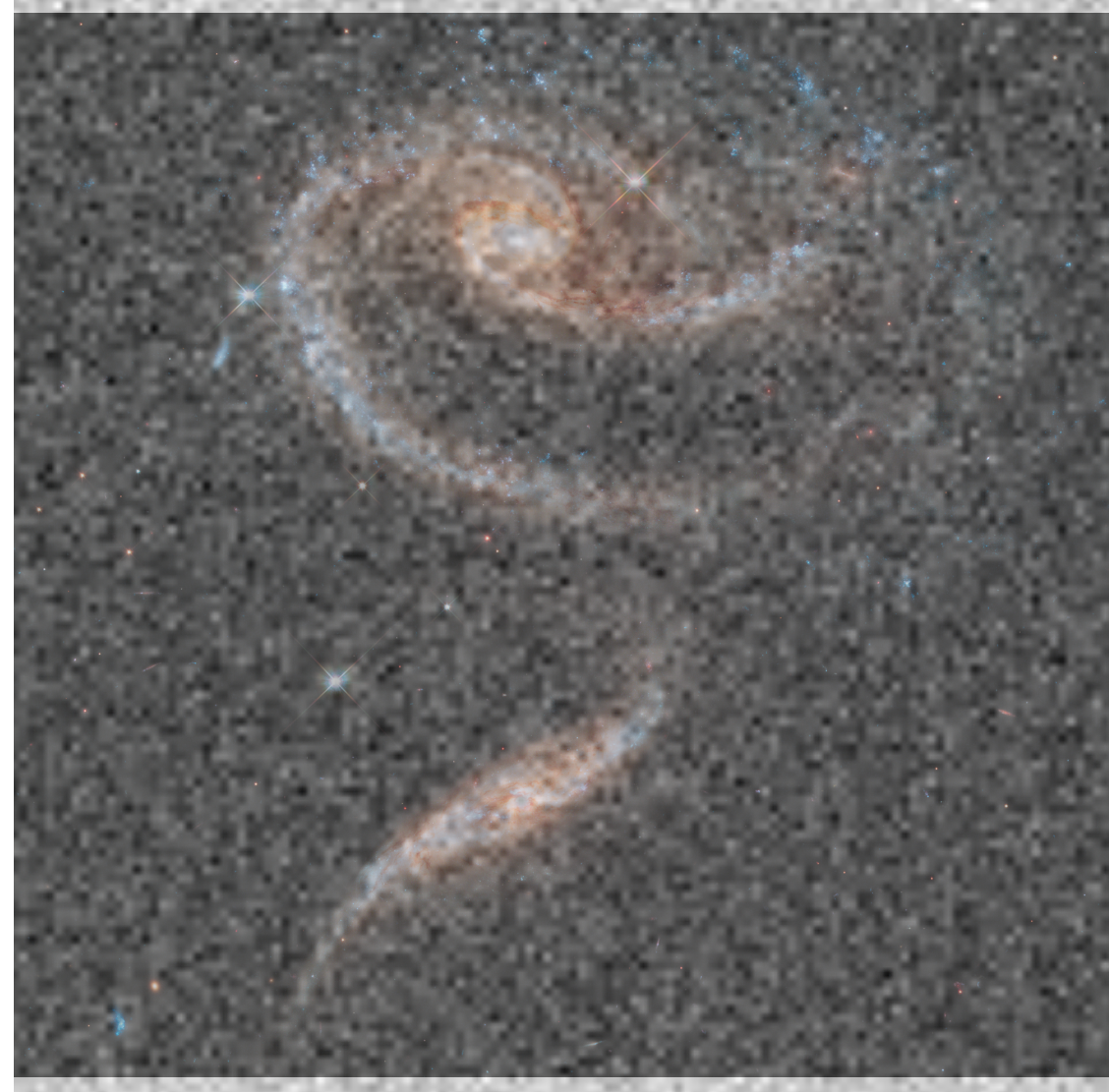


=

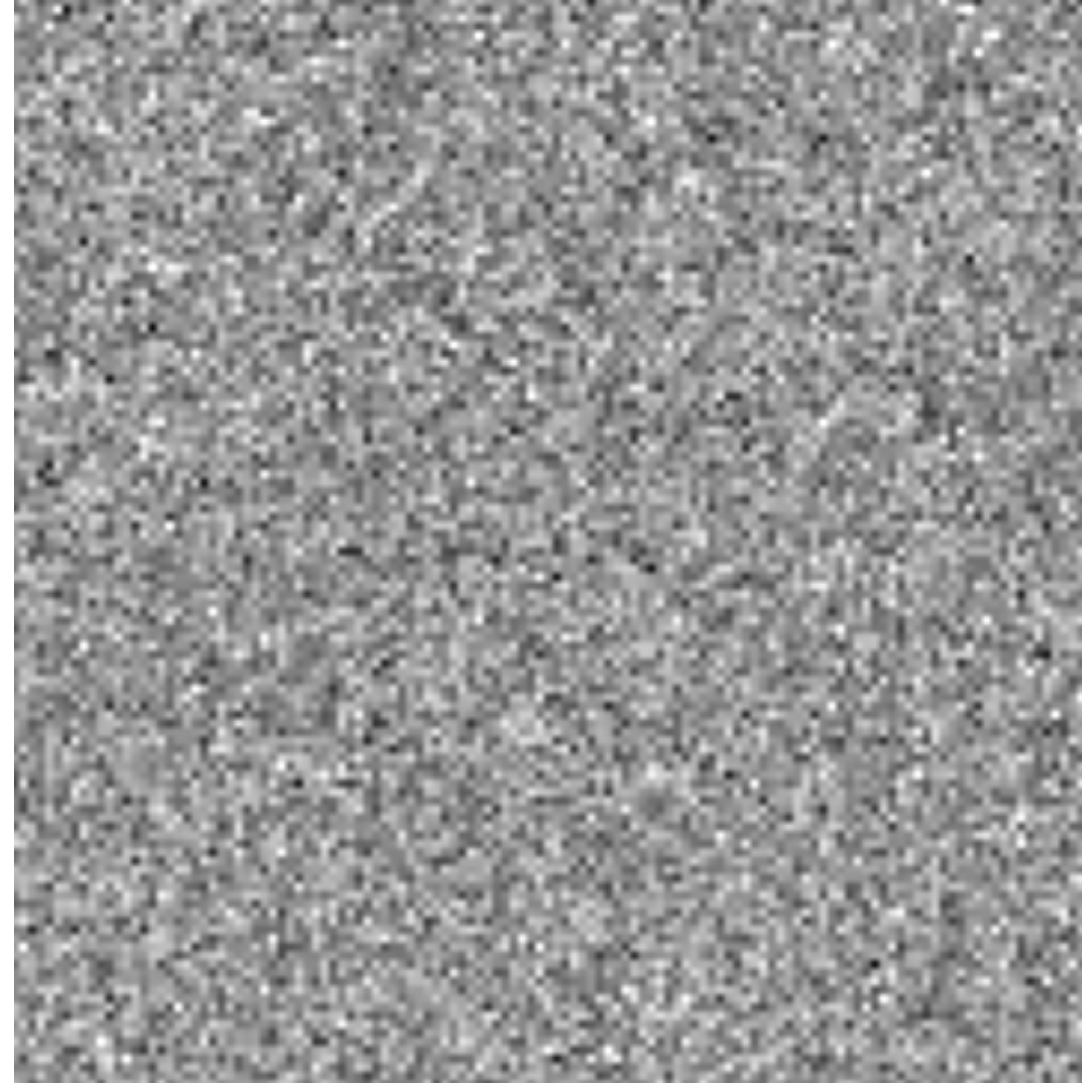


+

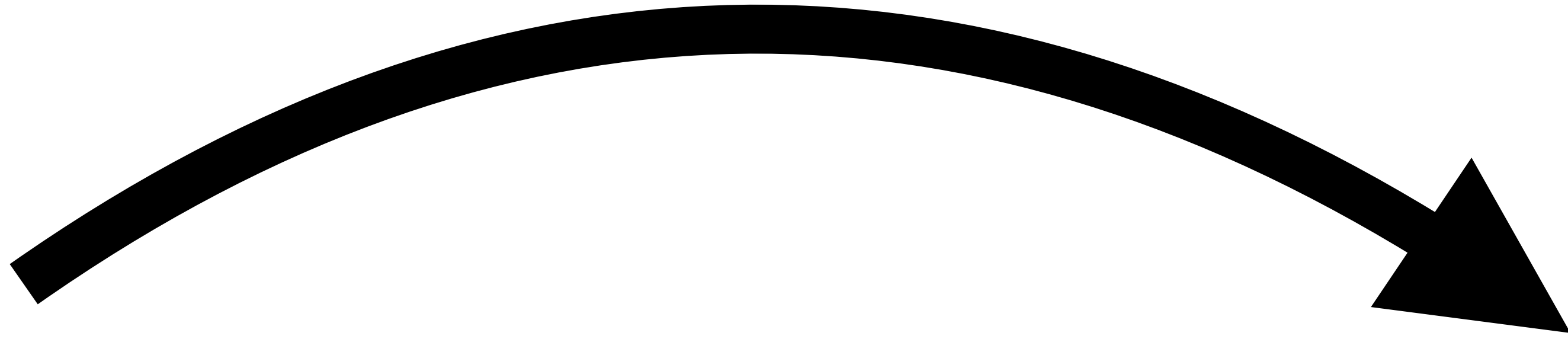


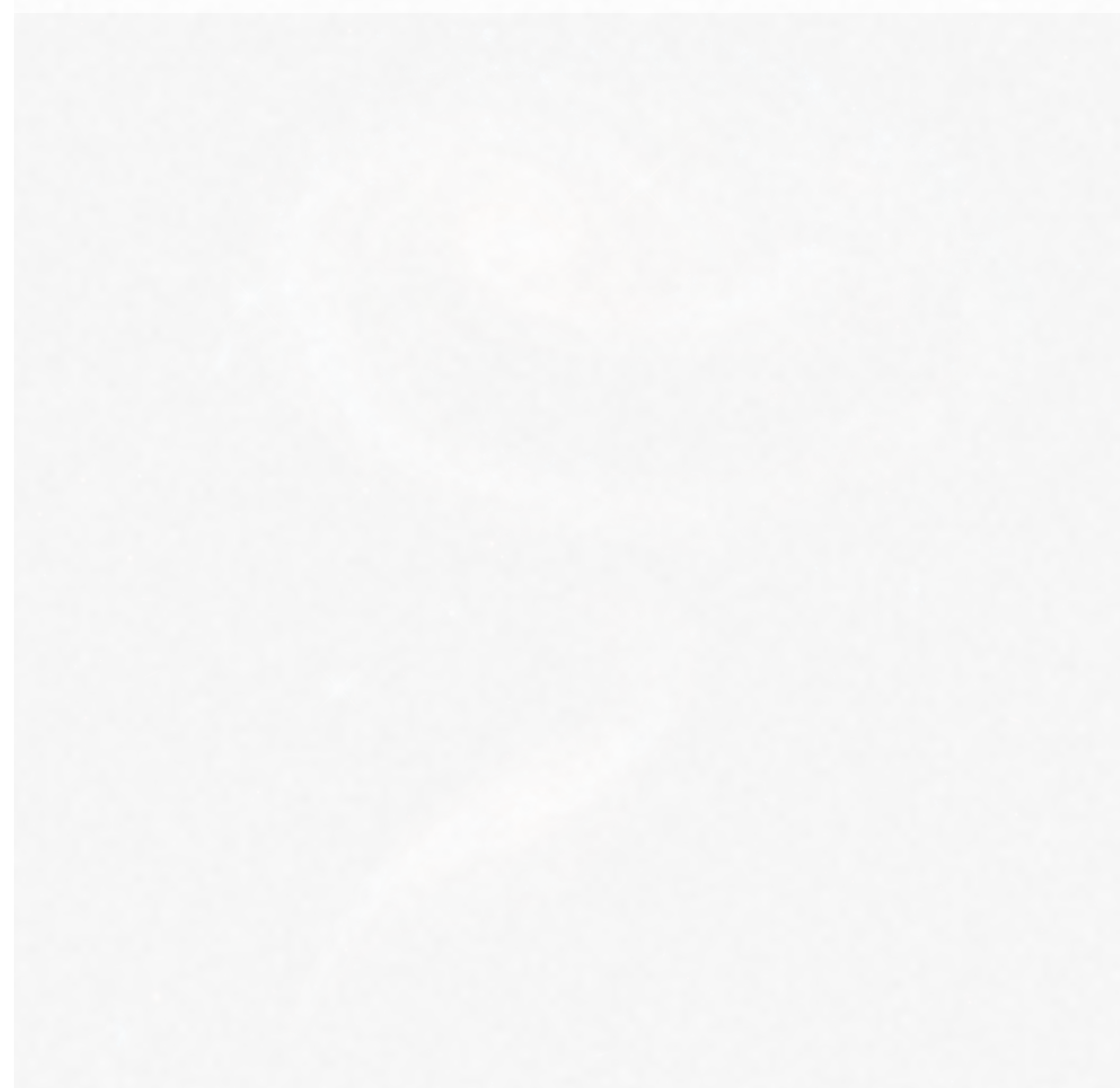


=

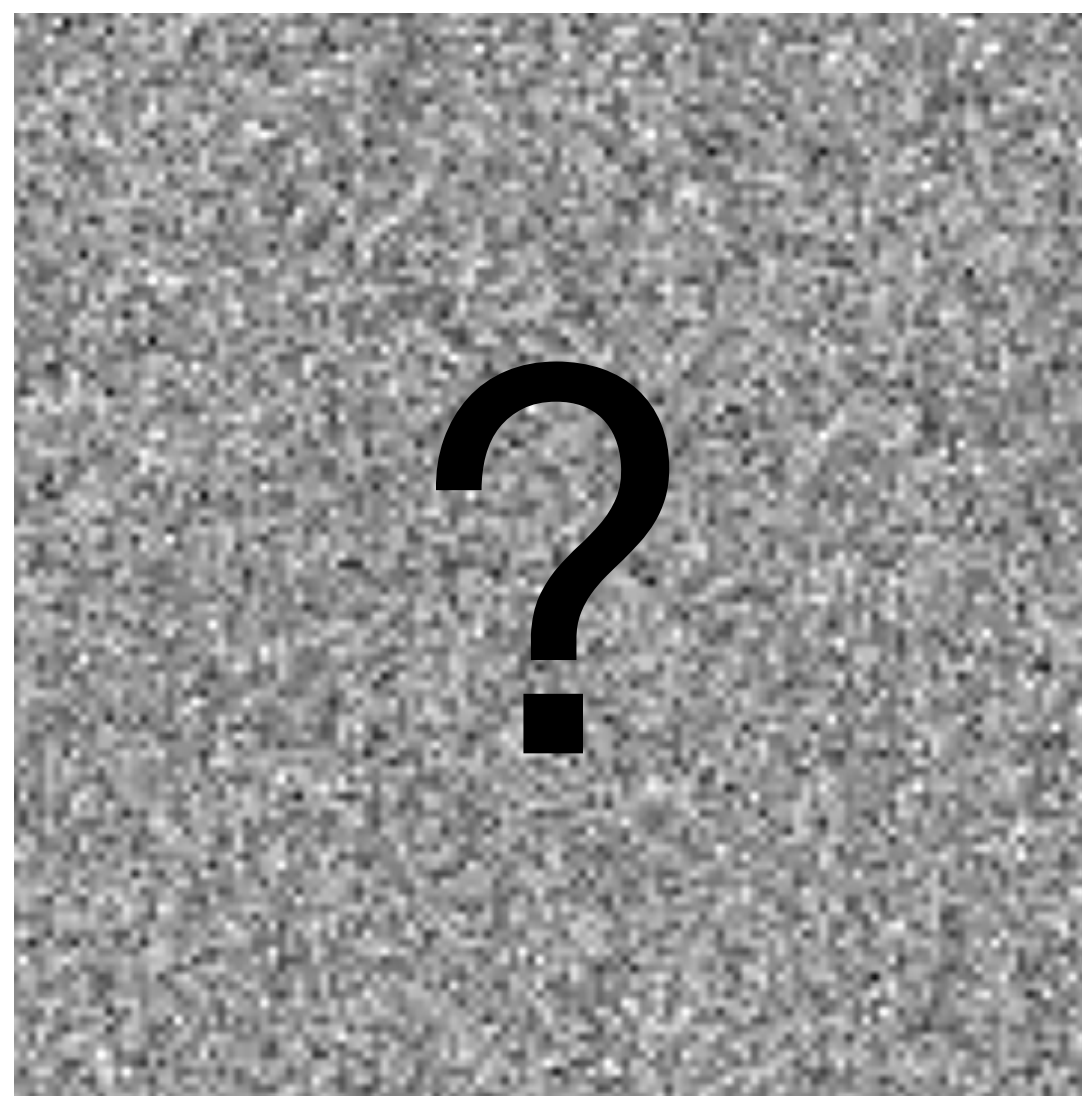


+

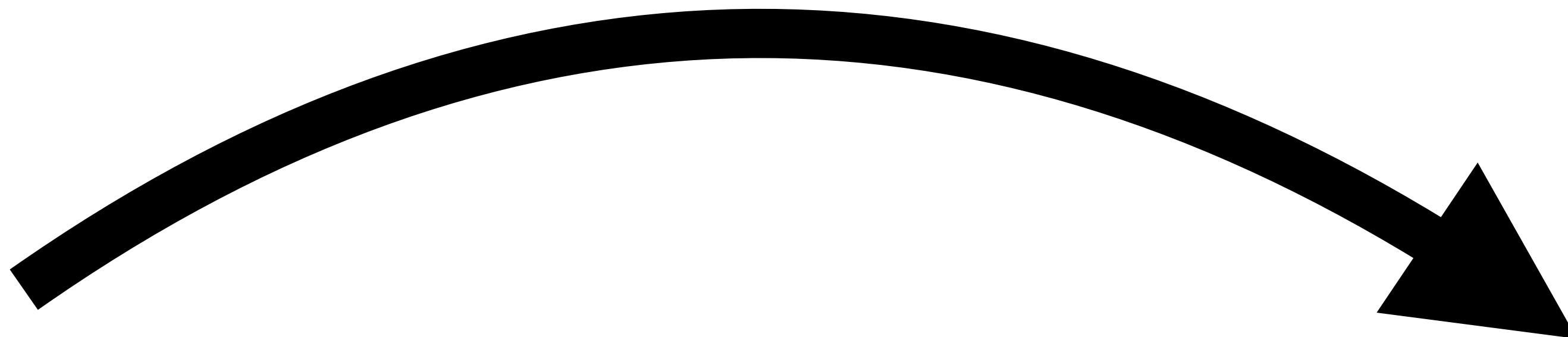




=

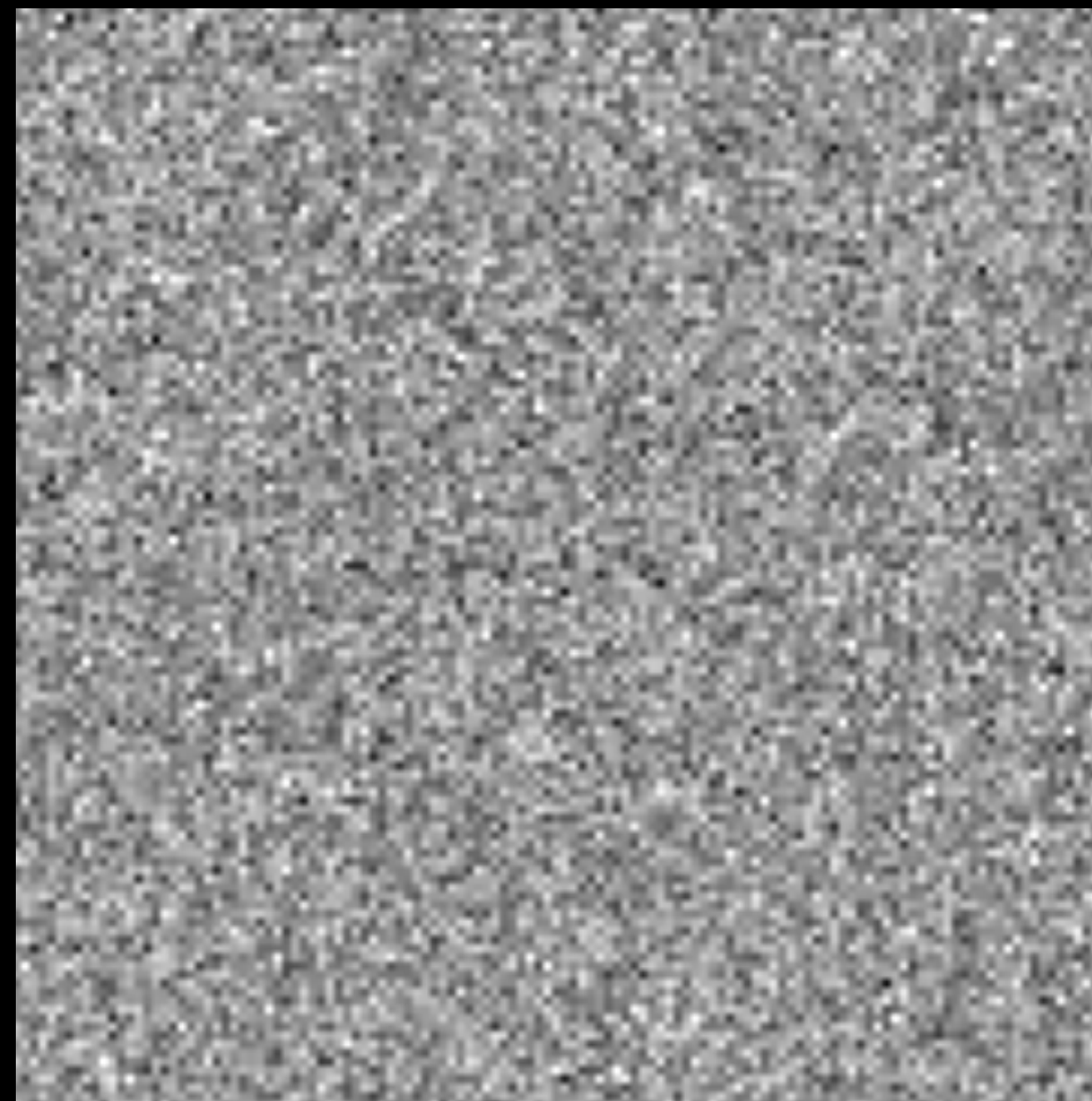


+

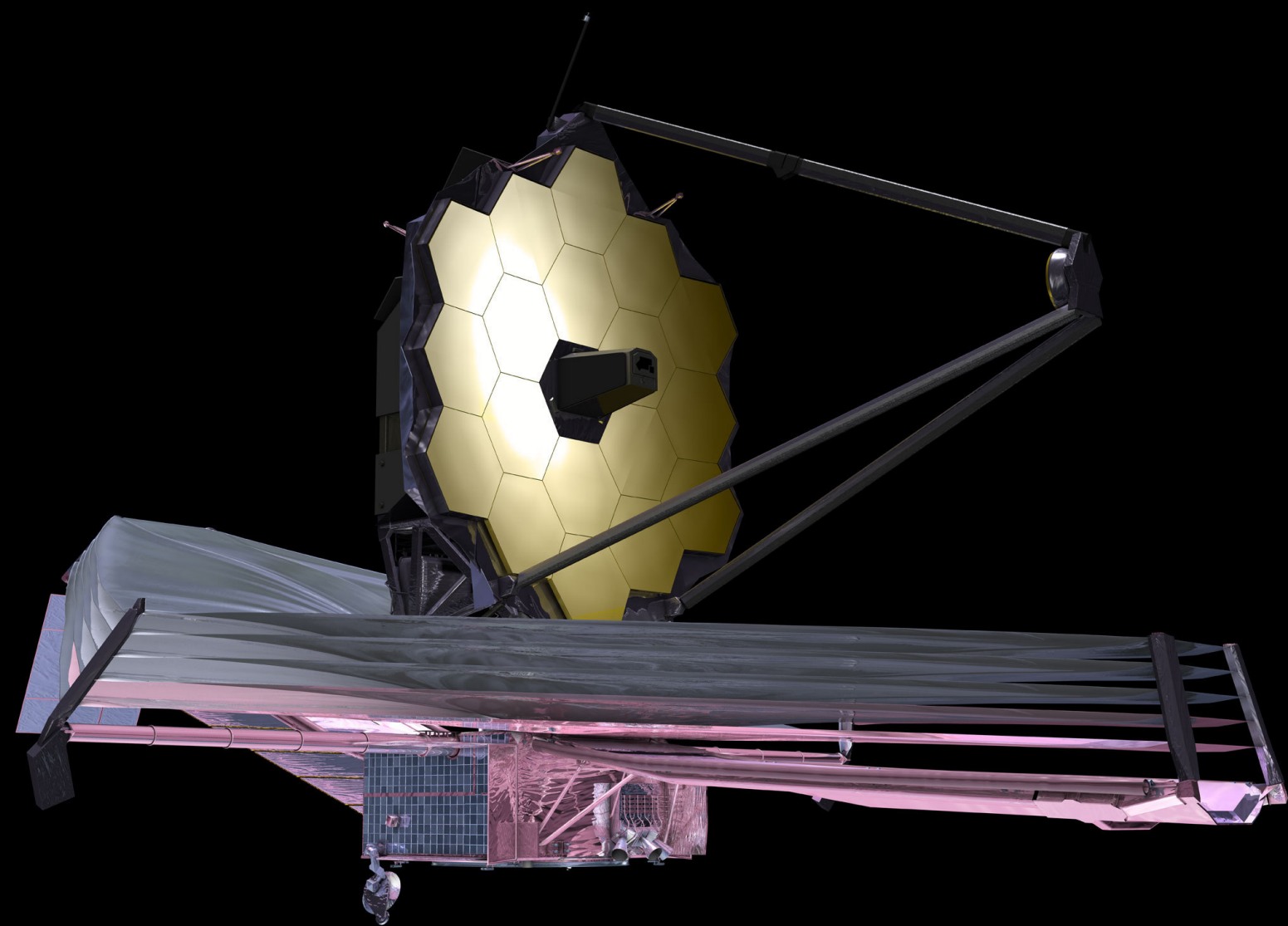


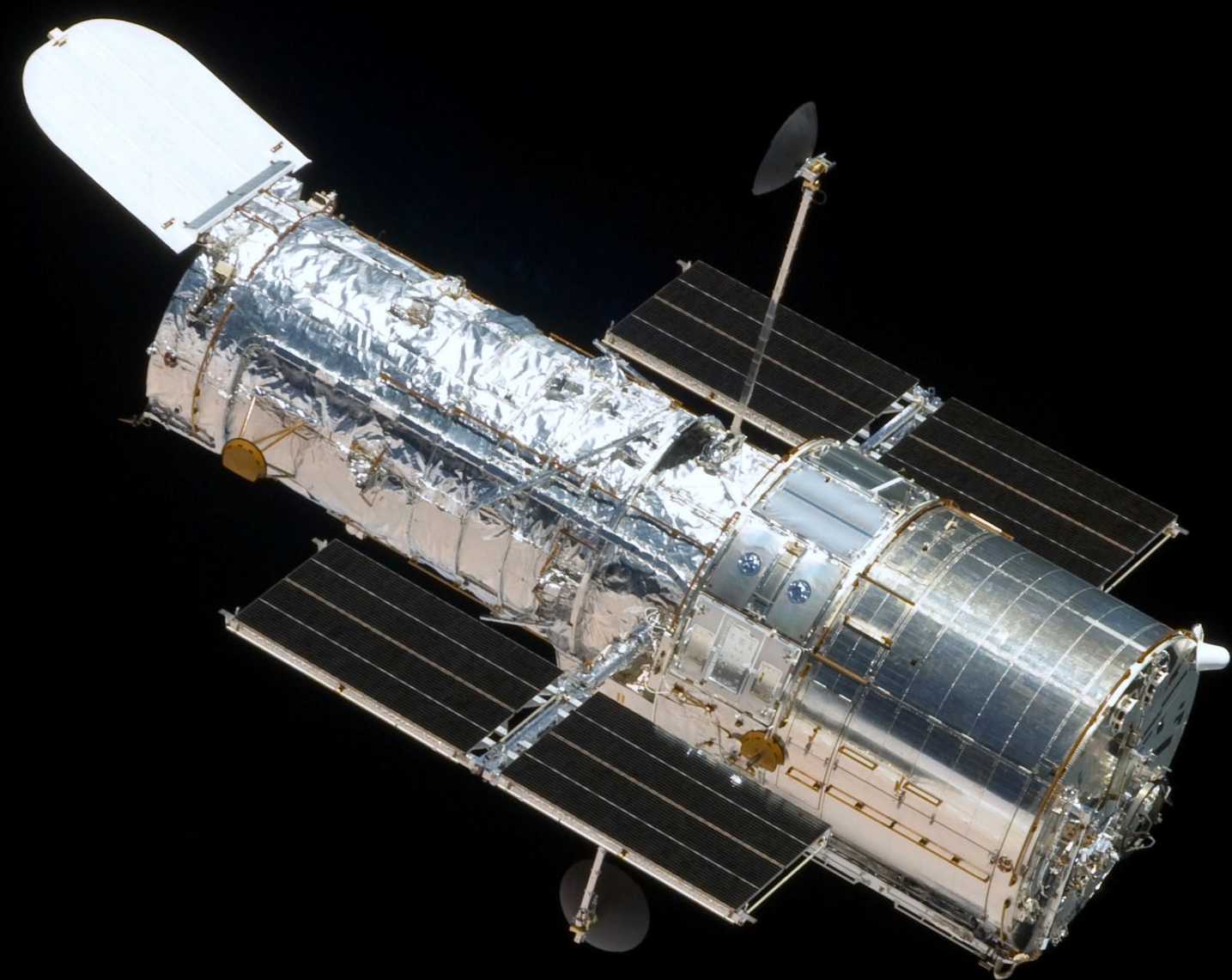
# Examples

Gaussian Noise



JWST Noise

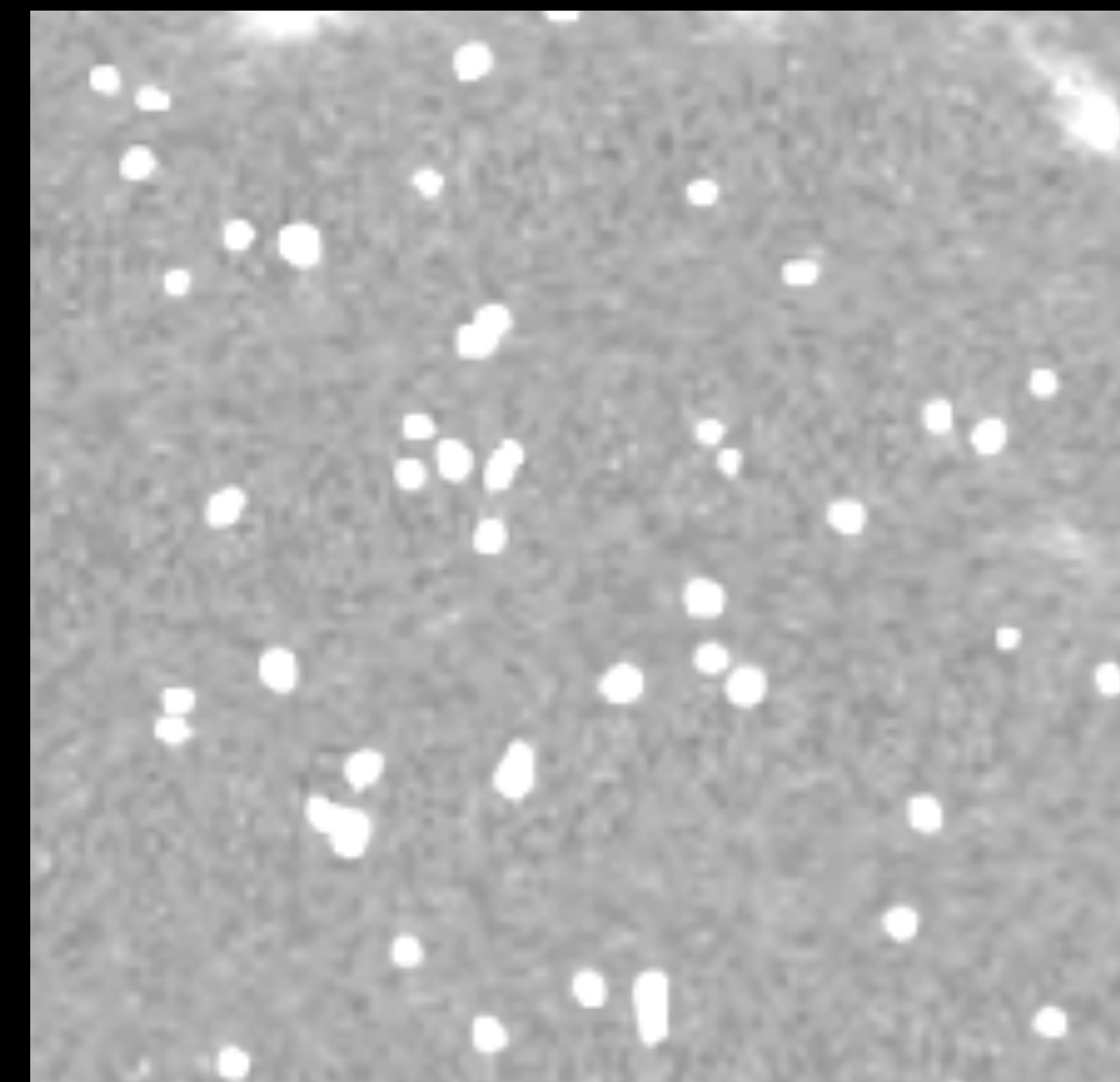




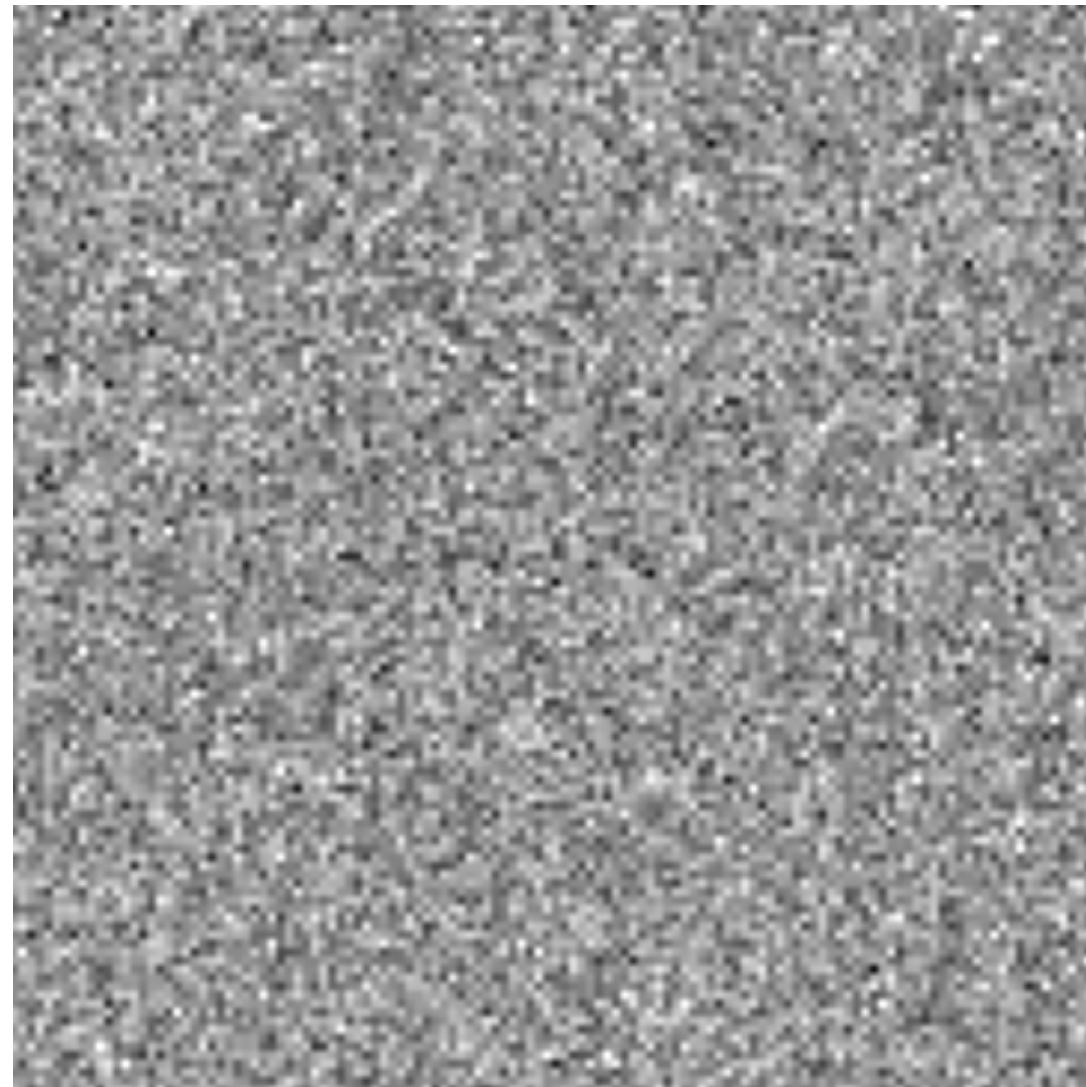
Gaussian Noise



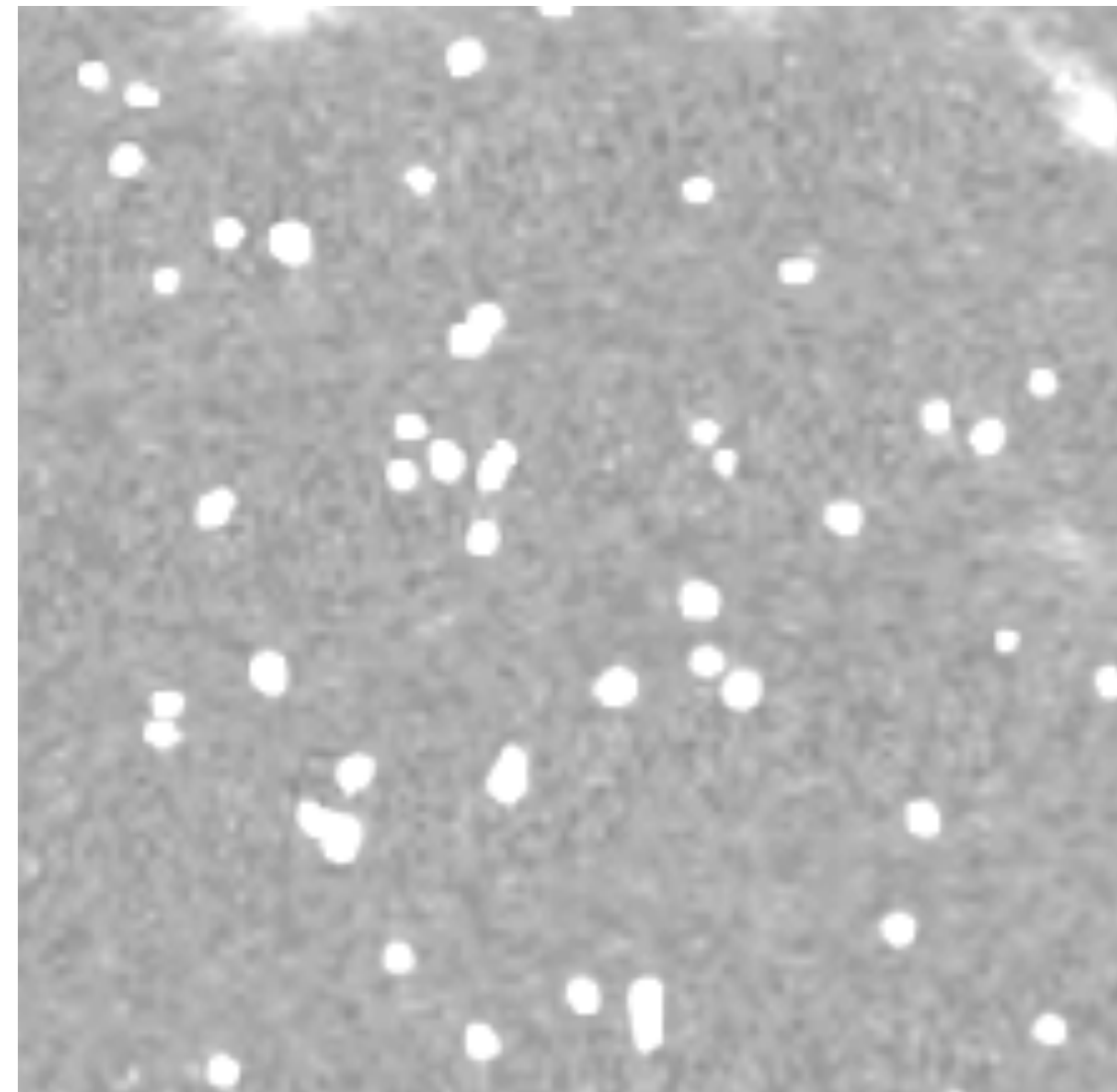
HST Noise



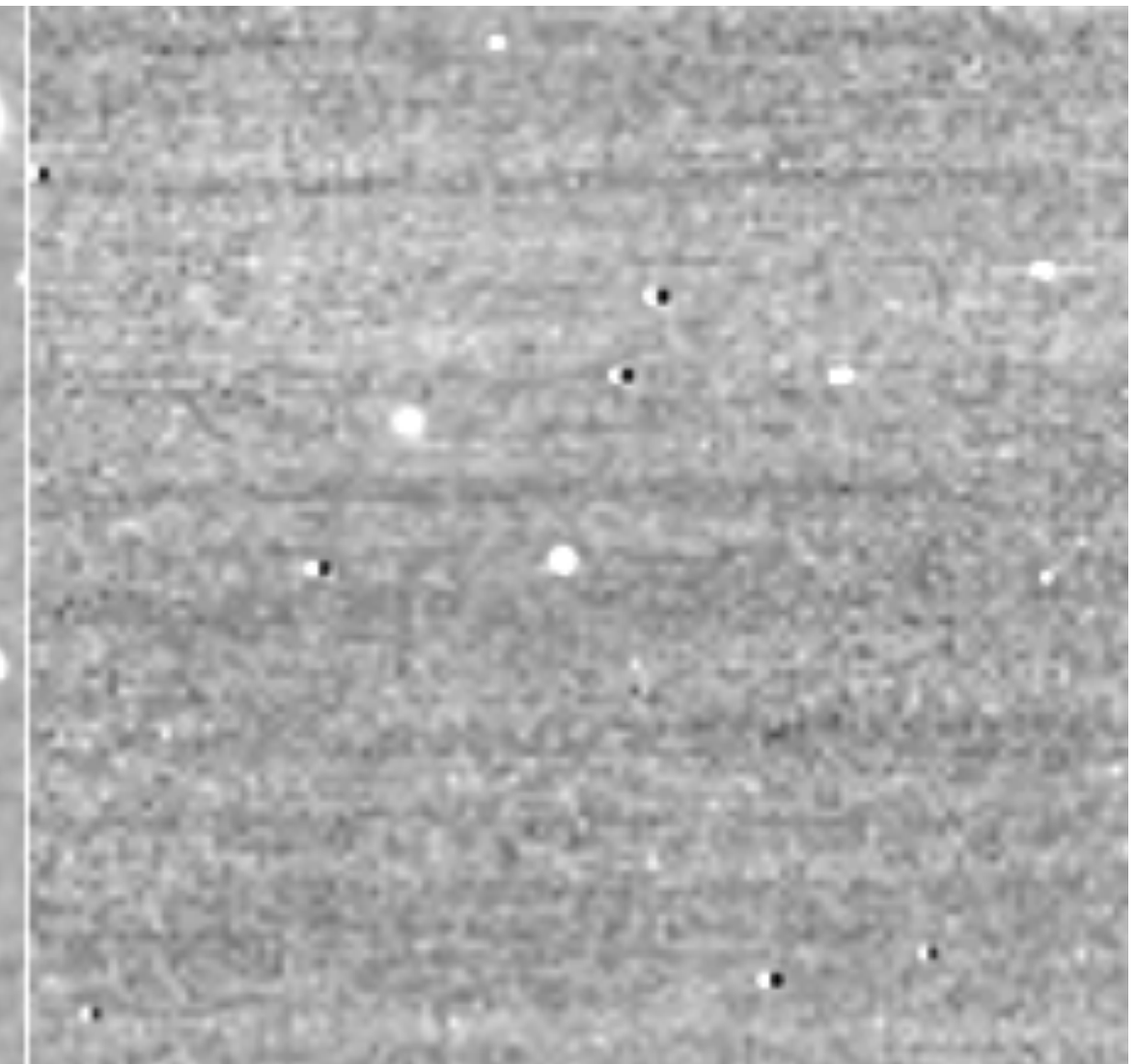
Gaussian Noise



HST noise



JWST noise



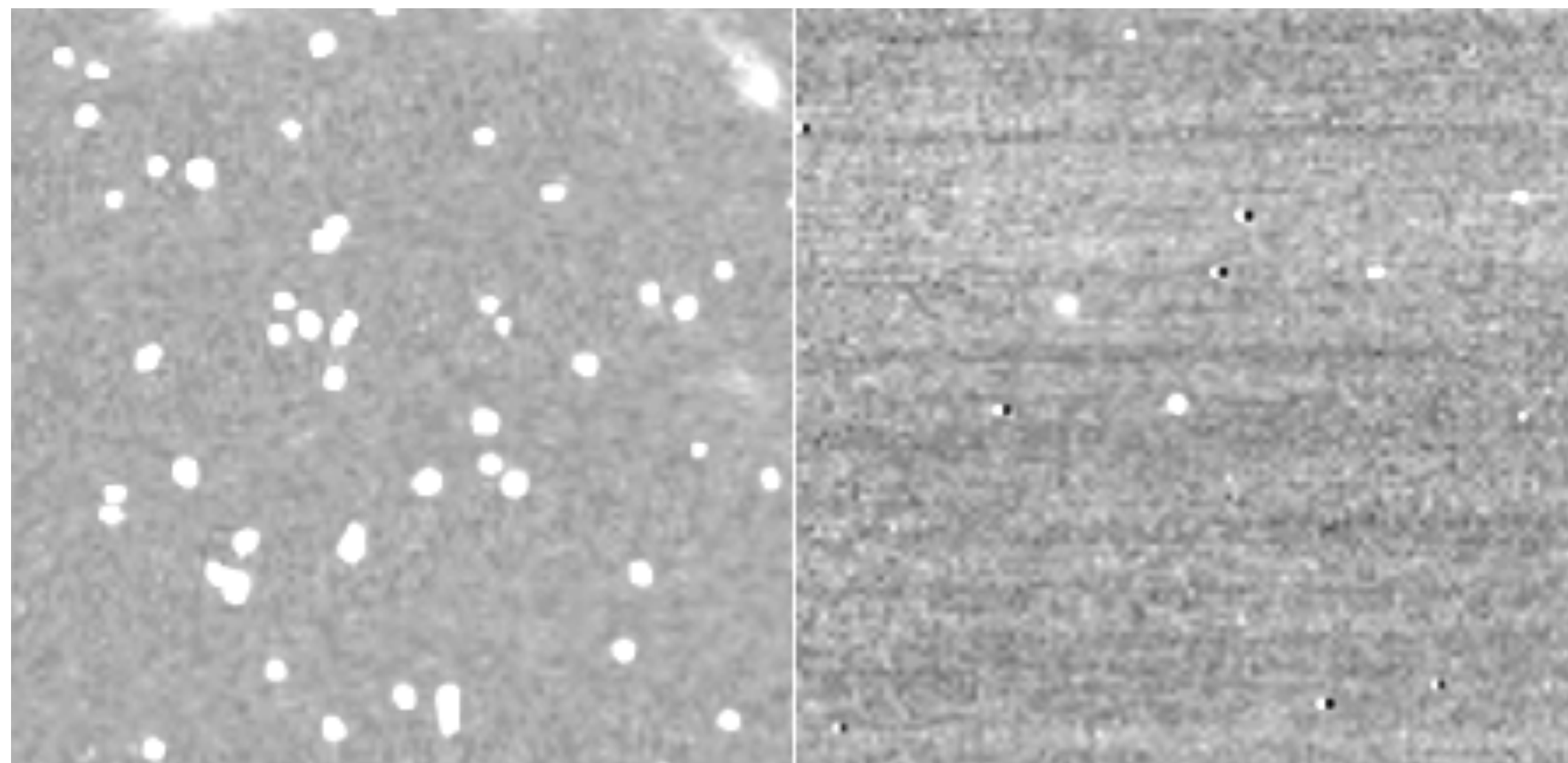
$$P(d|\theta) = \frac{1}{\sqrt{2\pi\sigma^2}} \exp - \frac{1}{2} \left( \frac{d - \mu(\theta)}{\sigma} \right)^2$$

?

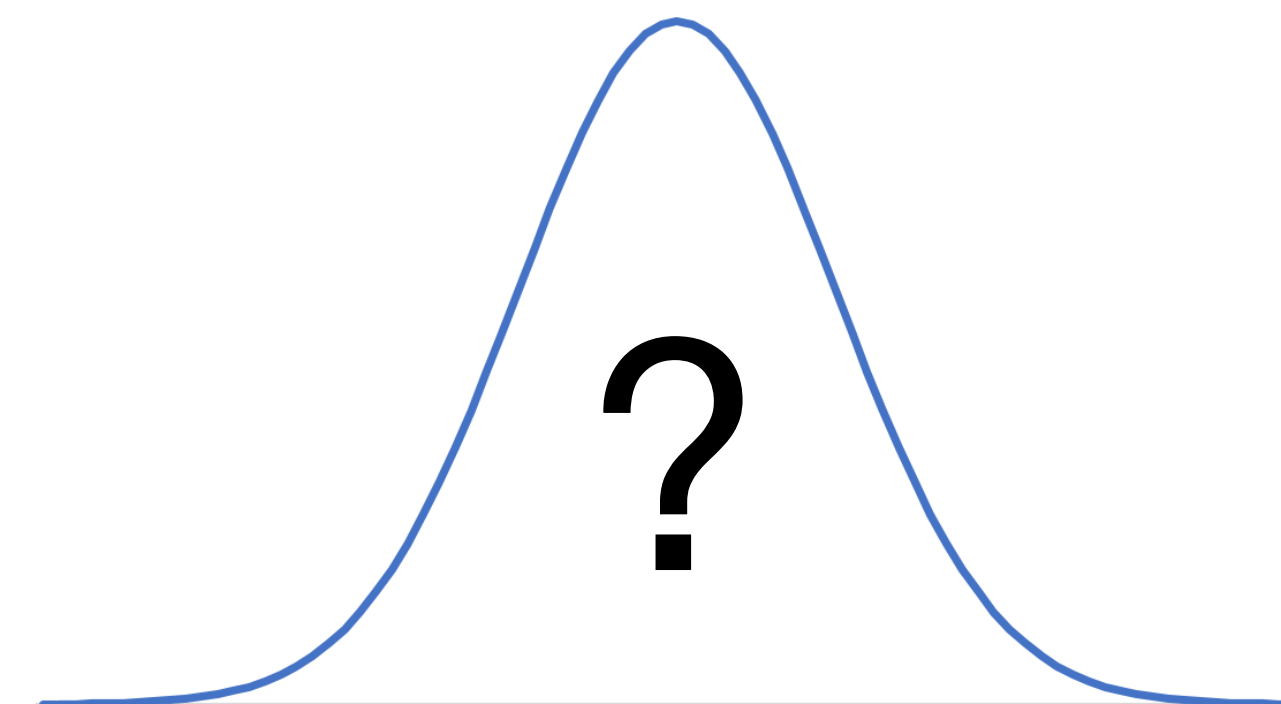
?



# Learn Noise Distribution



$Q(x)$



**Alternative?**

$Q(x)$

# Alternative?

- Instead learn  $\nabla_x \log Q(x)$ .

~~$Q(x)$~~



$\nabla_x \log Q(x)$

# Score-based generative models



$$\nabla_x \log Q(x)$$

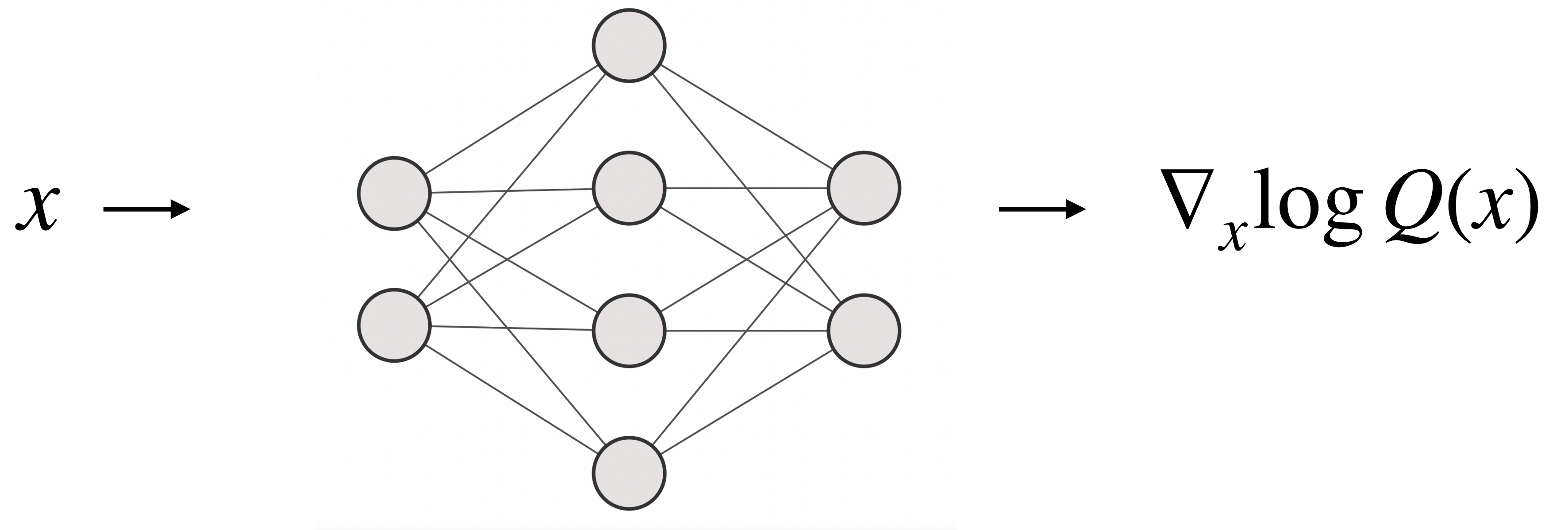
# Score-based generative models



$$\nabla_x \log Q(x)$$

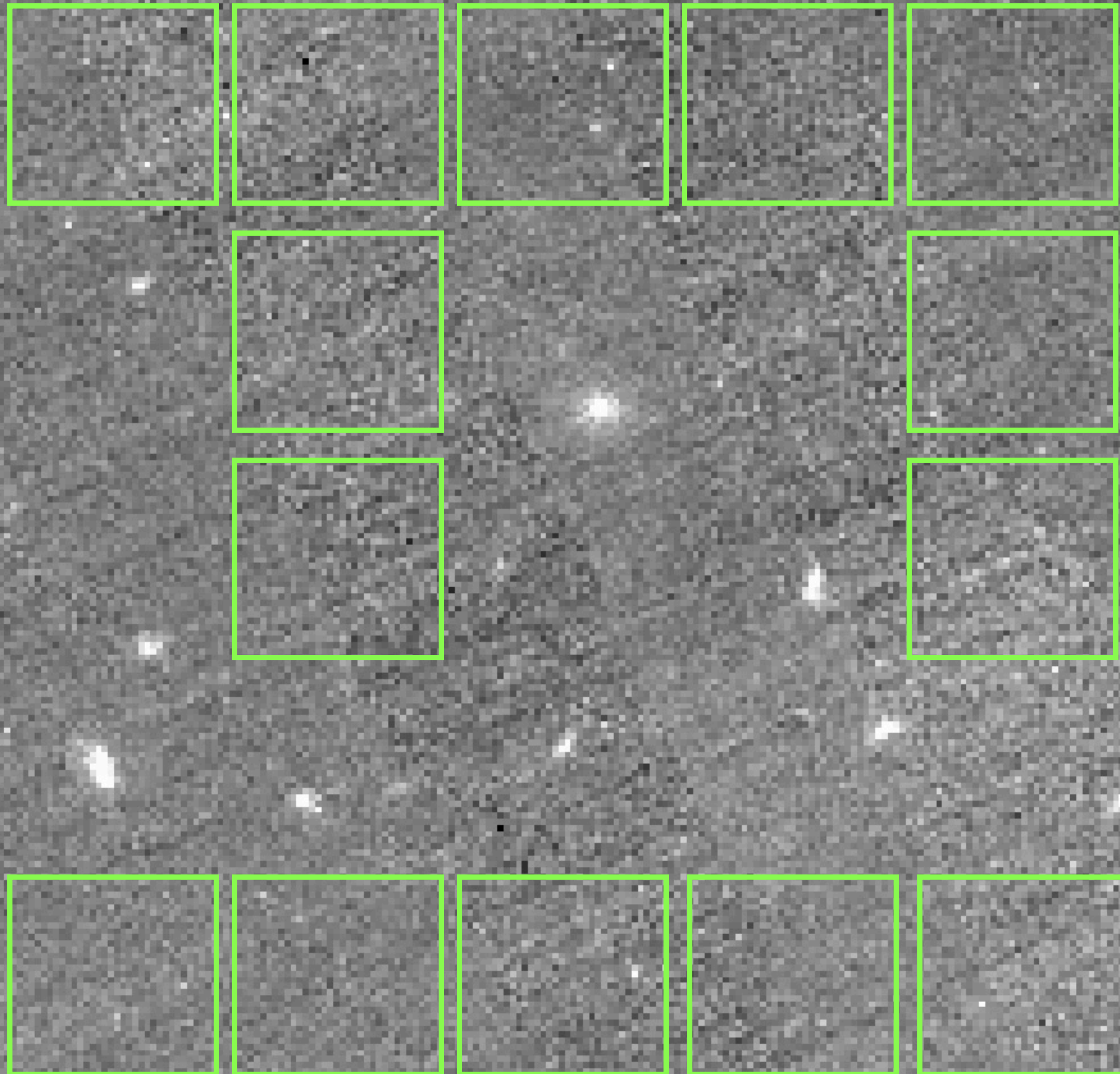
# The Machine Learning

- Train neural network to predict  $\nabla_x \log Q(x)$  using score matching techniques.



Machine Learning

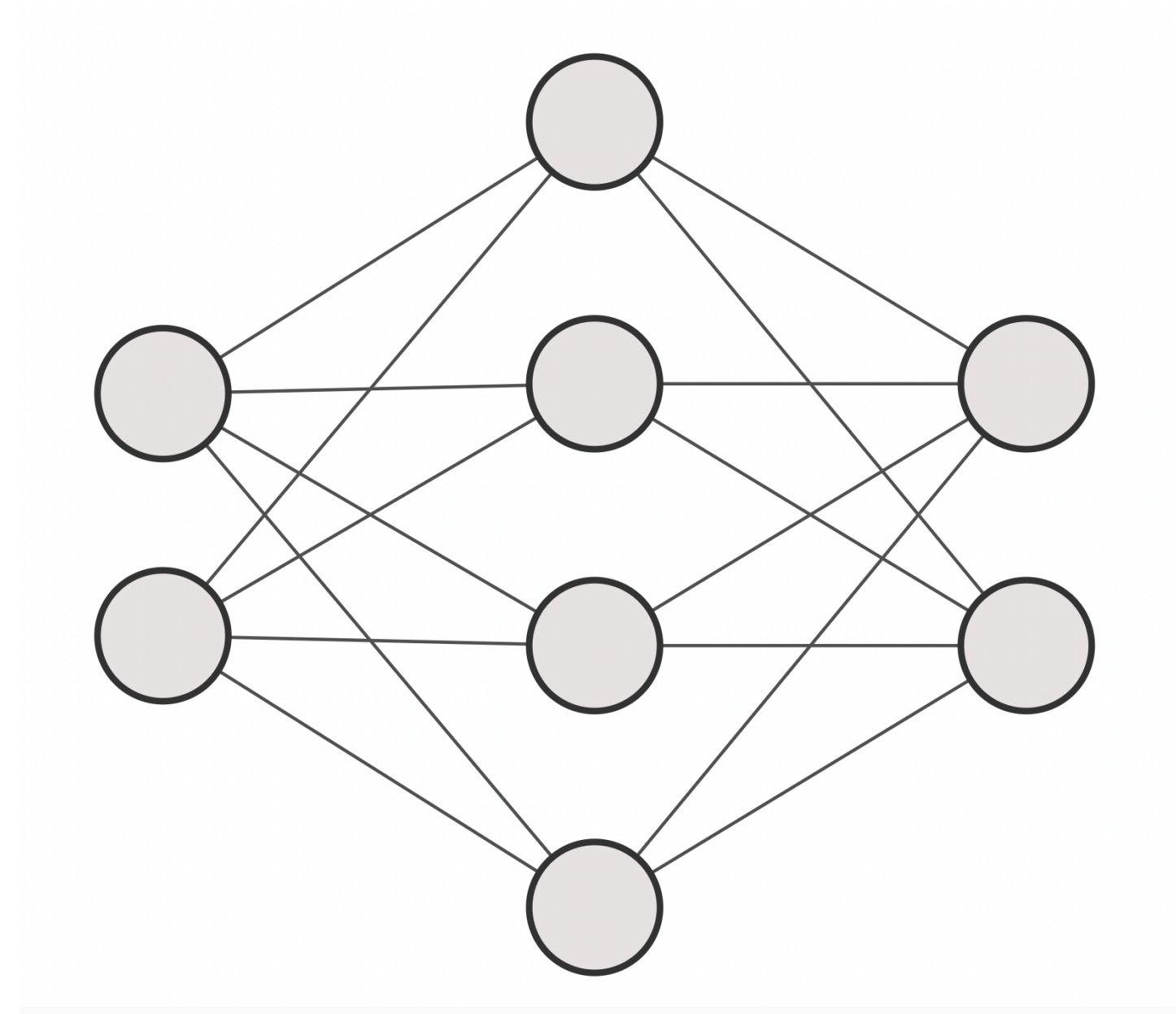
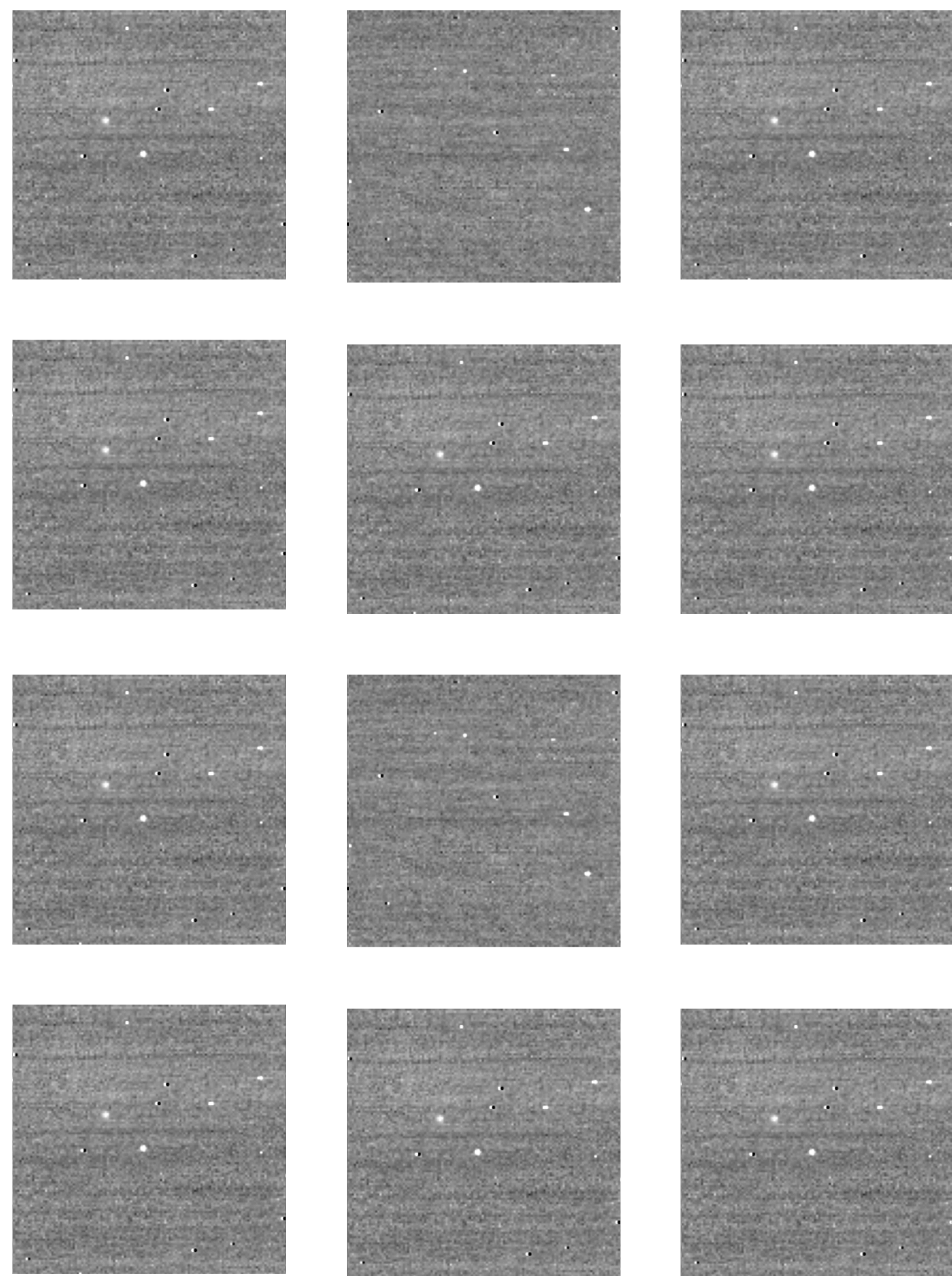






# The Machine Learning

- Train neural network to predict  $\nabla_x \log Q(x)$  using score matching techniques.



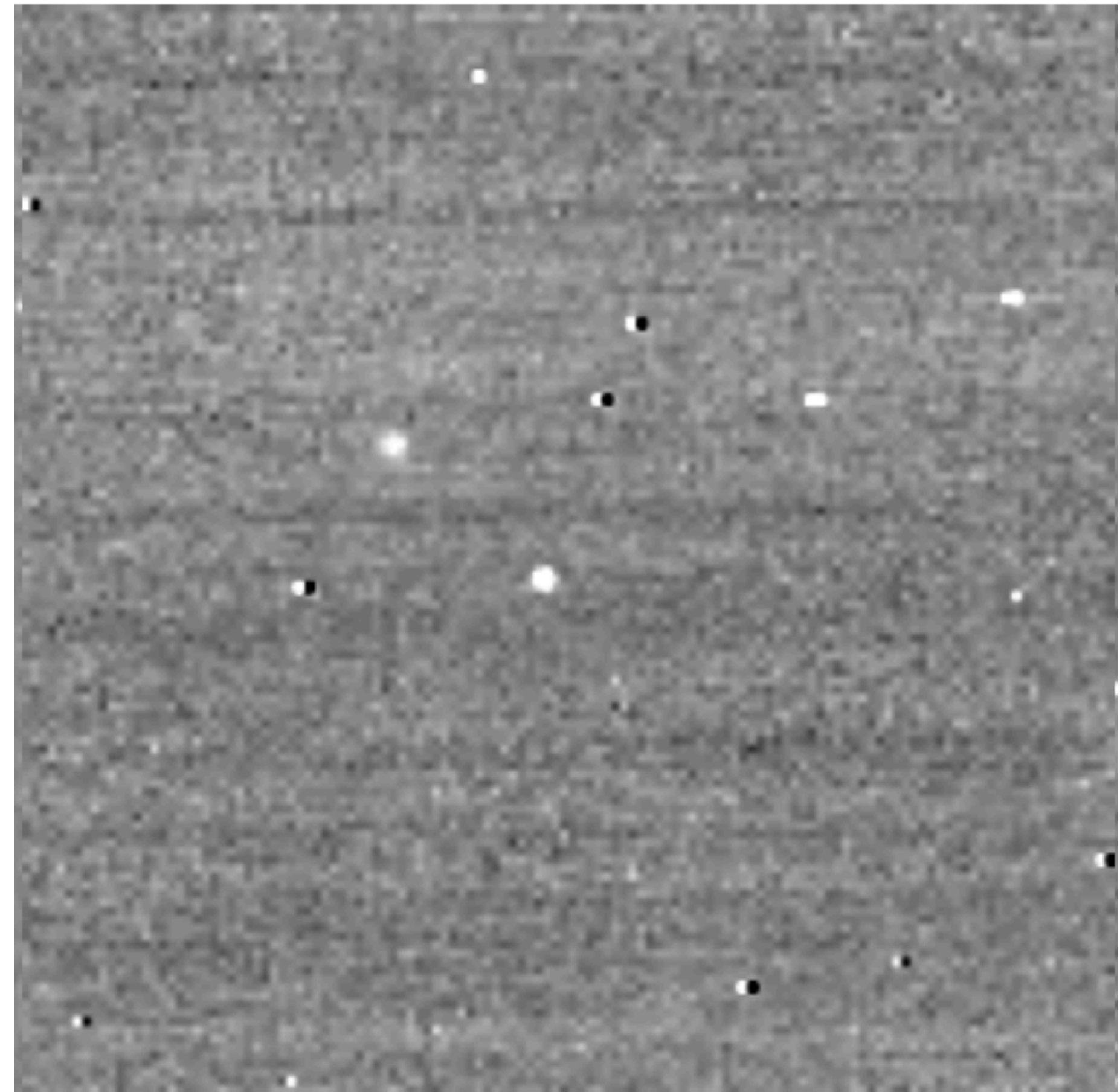
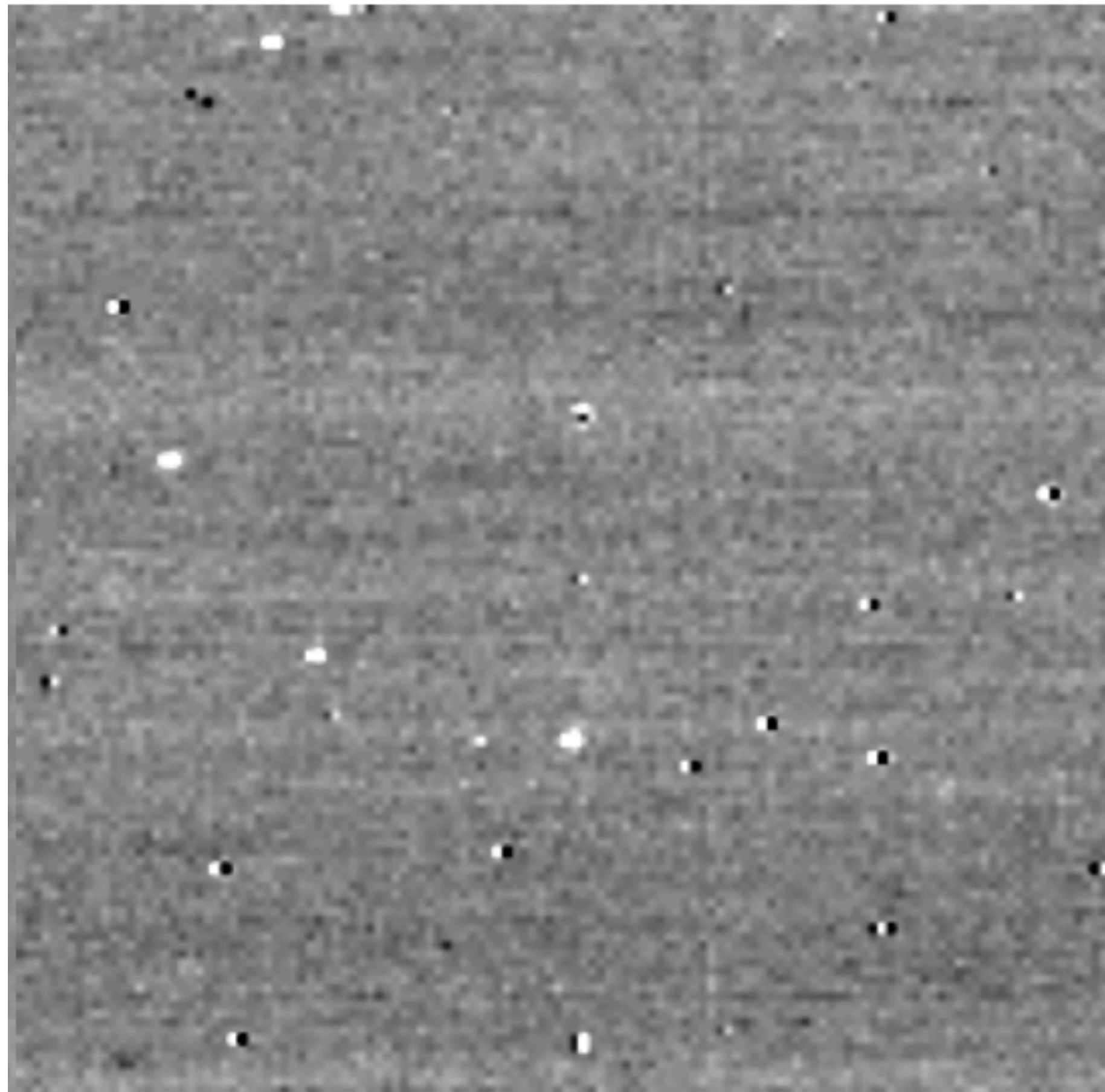
$$\nabla_x \log Q(x)$$

Machine Learning

# Results

# Noise Generation

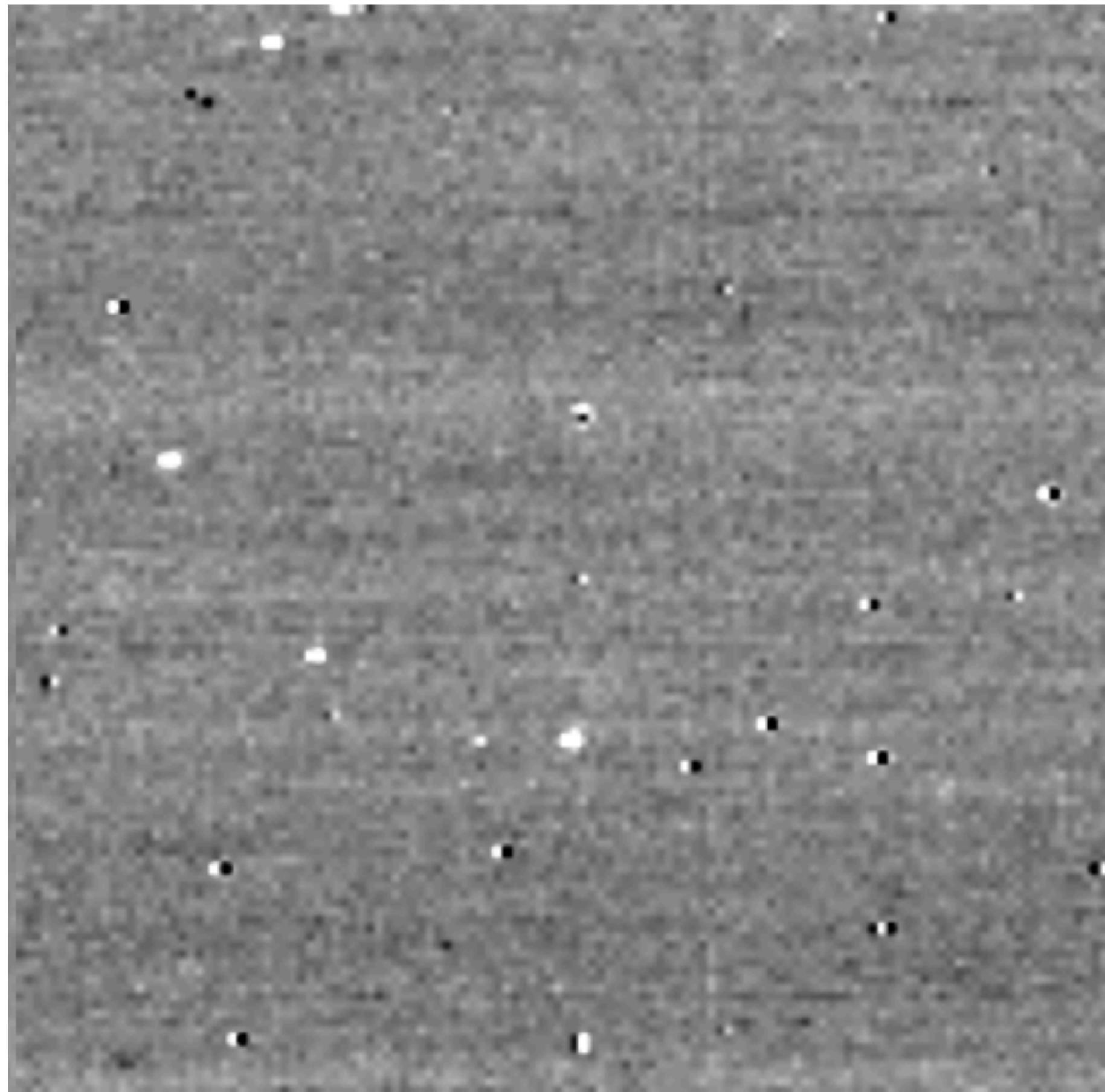
- Can sample new noise. Which one is real?



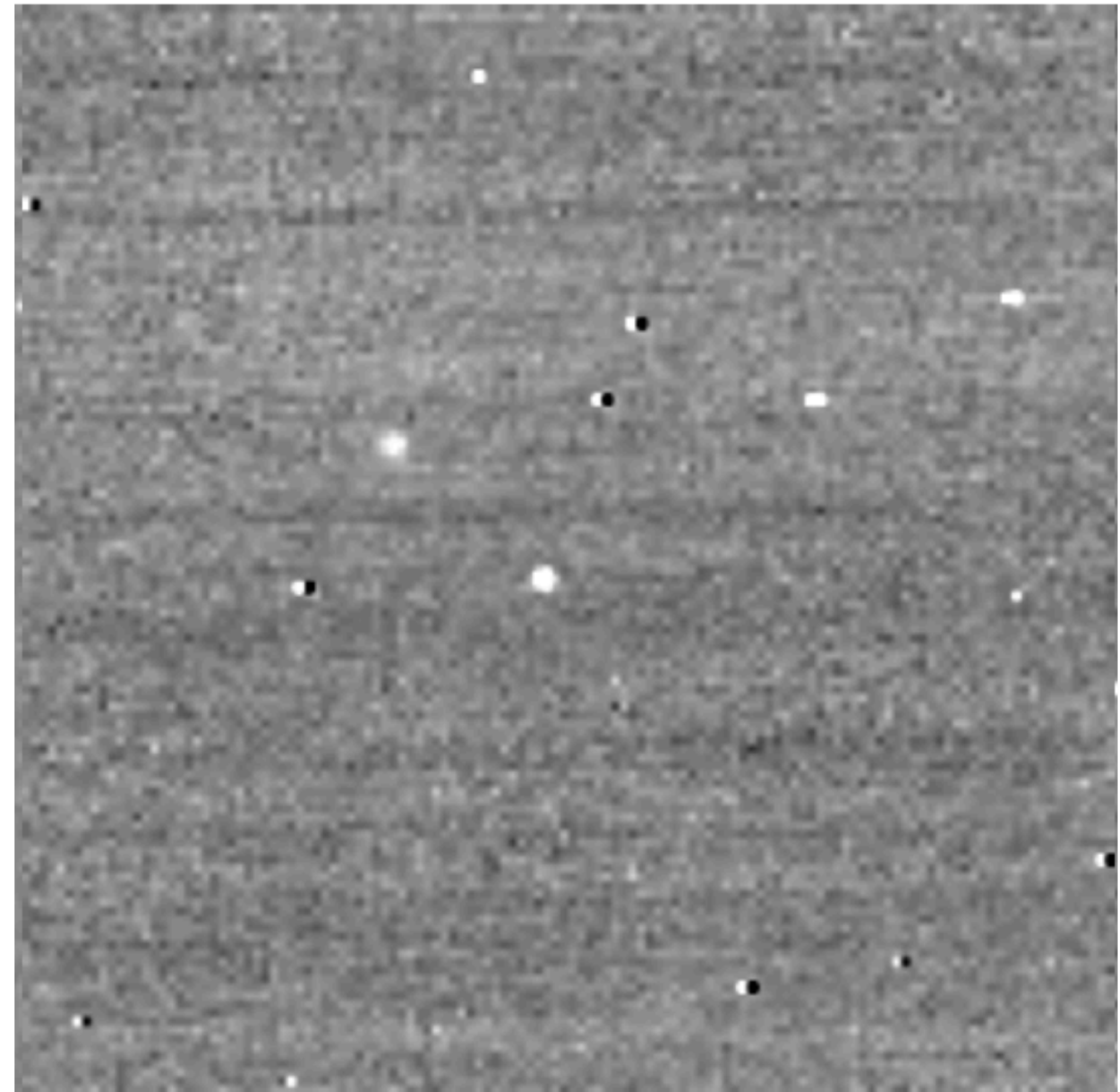
# Noise Generation

- Can sample new noise. Which one is real?

Real

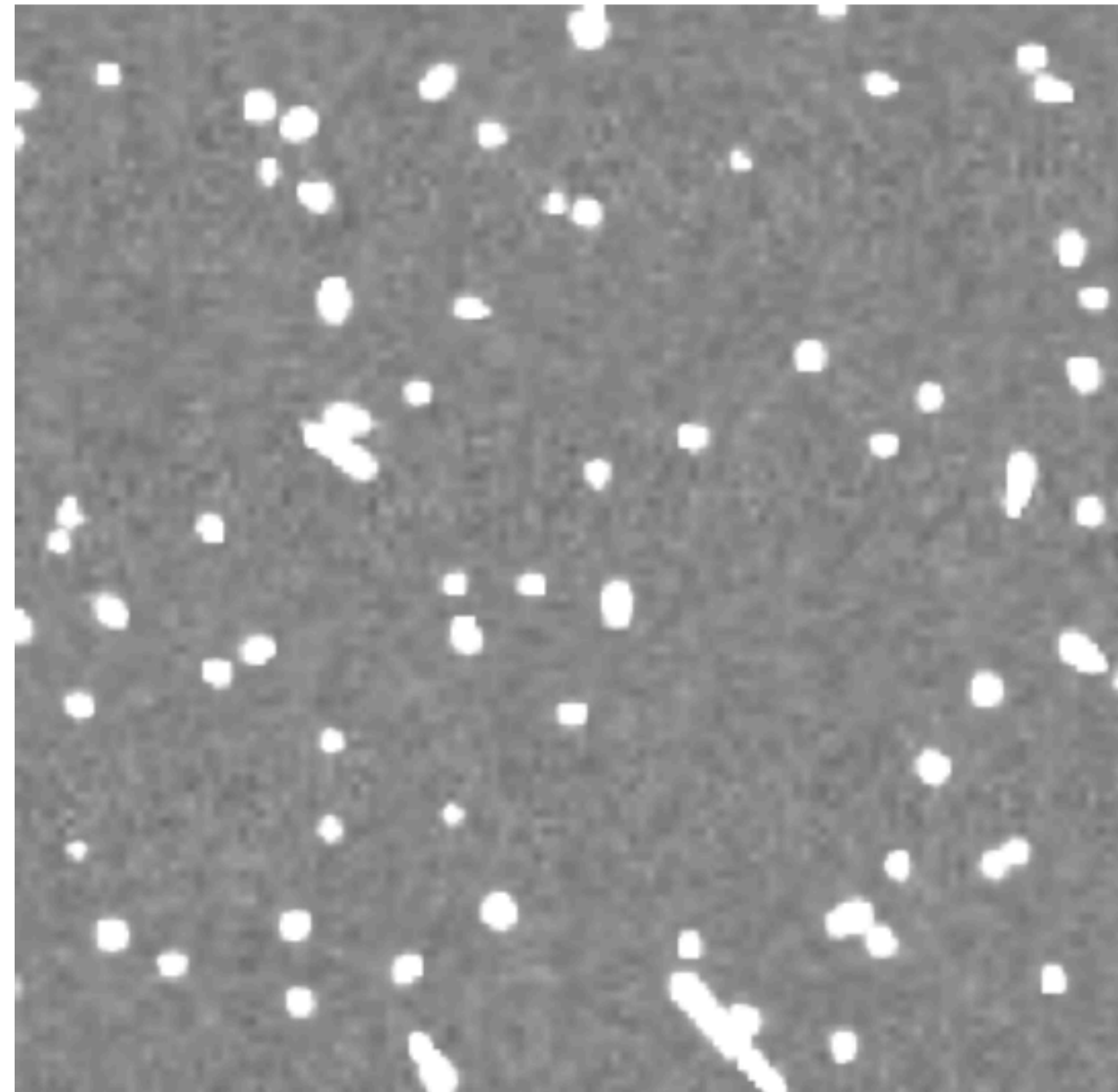
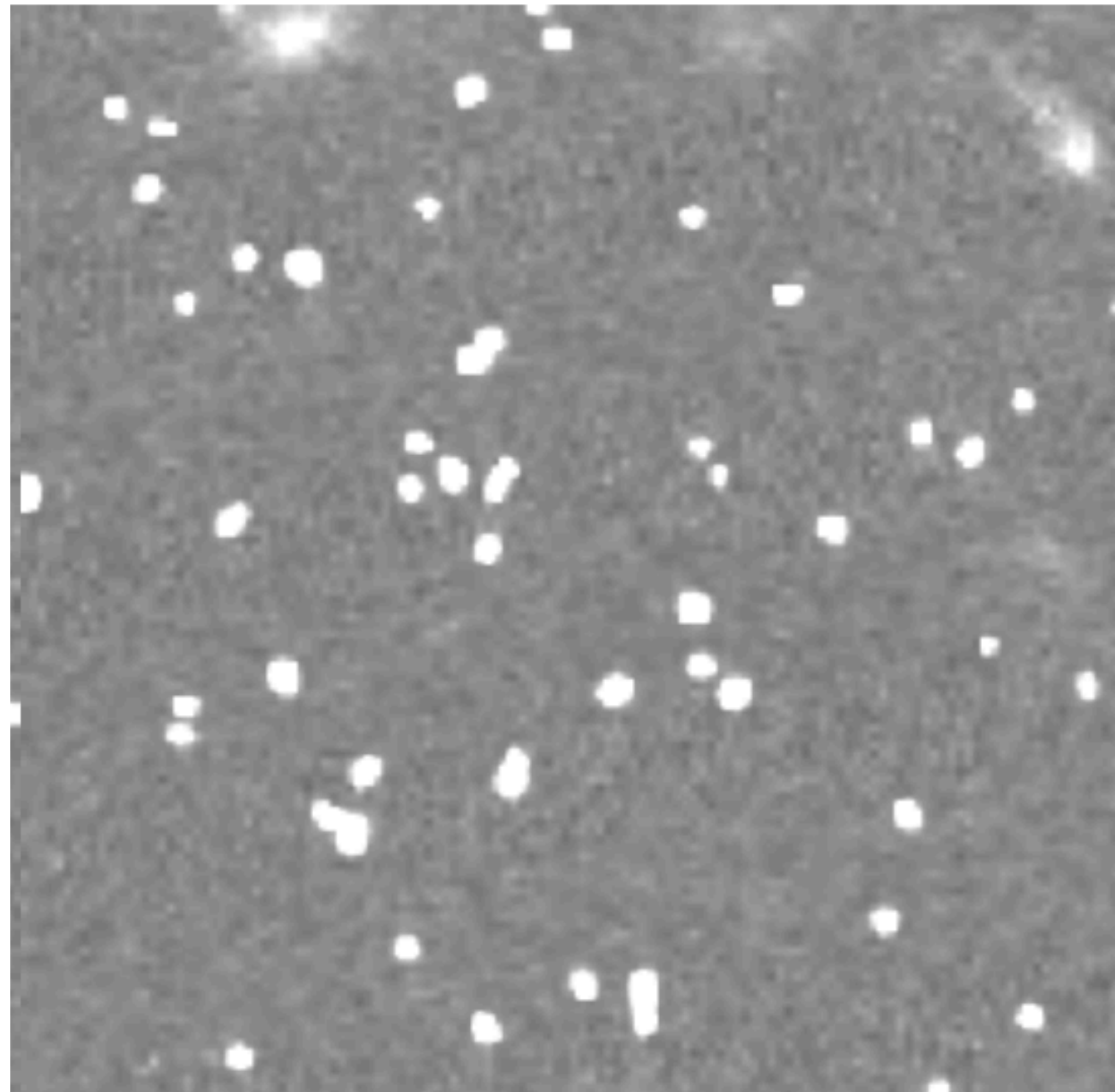


Fake (generated)



# Noise Generation

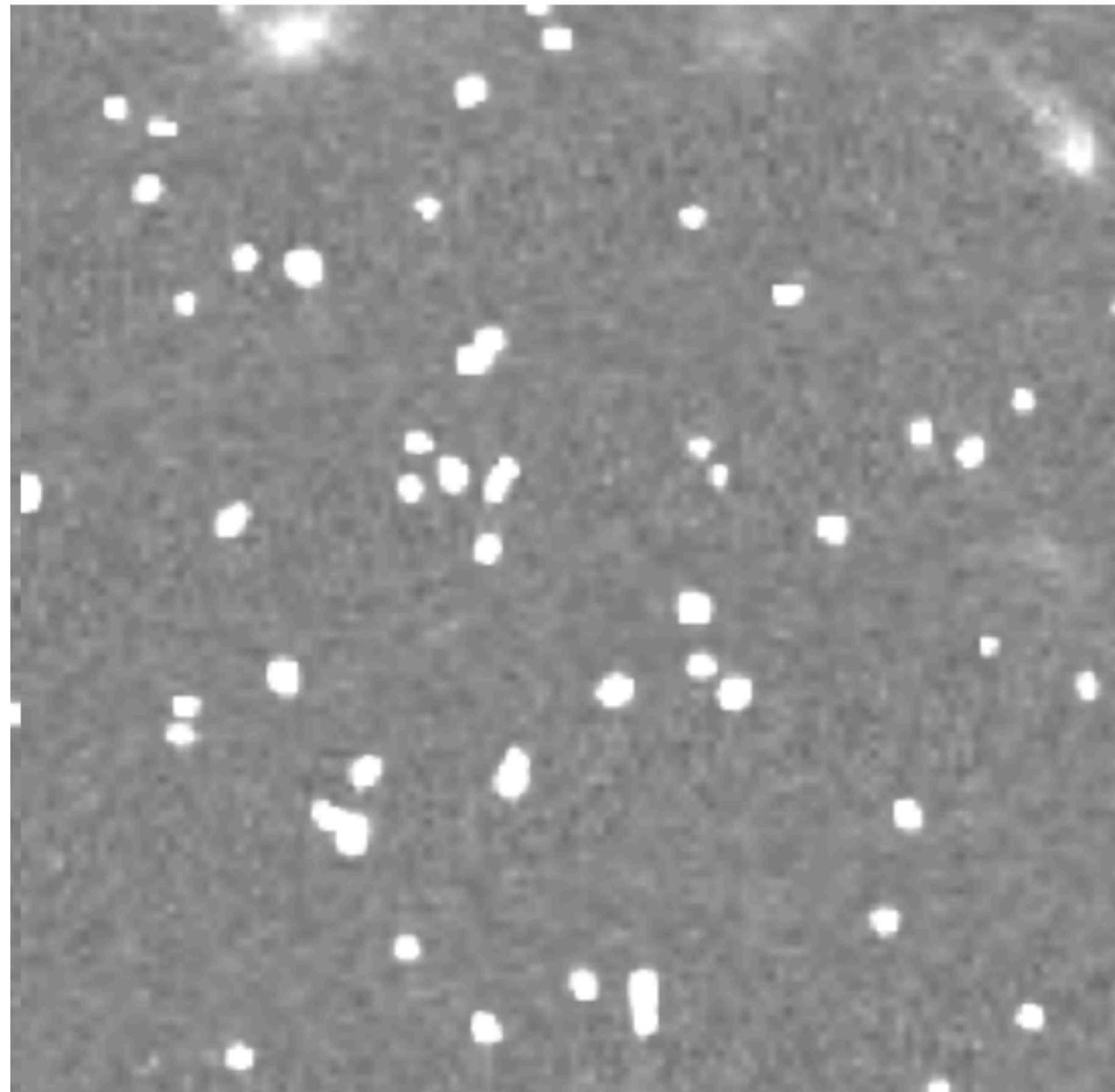
- Can sample new noise. Which one is real?



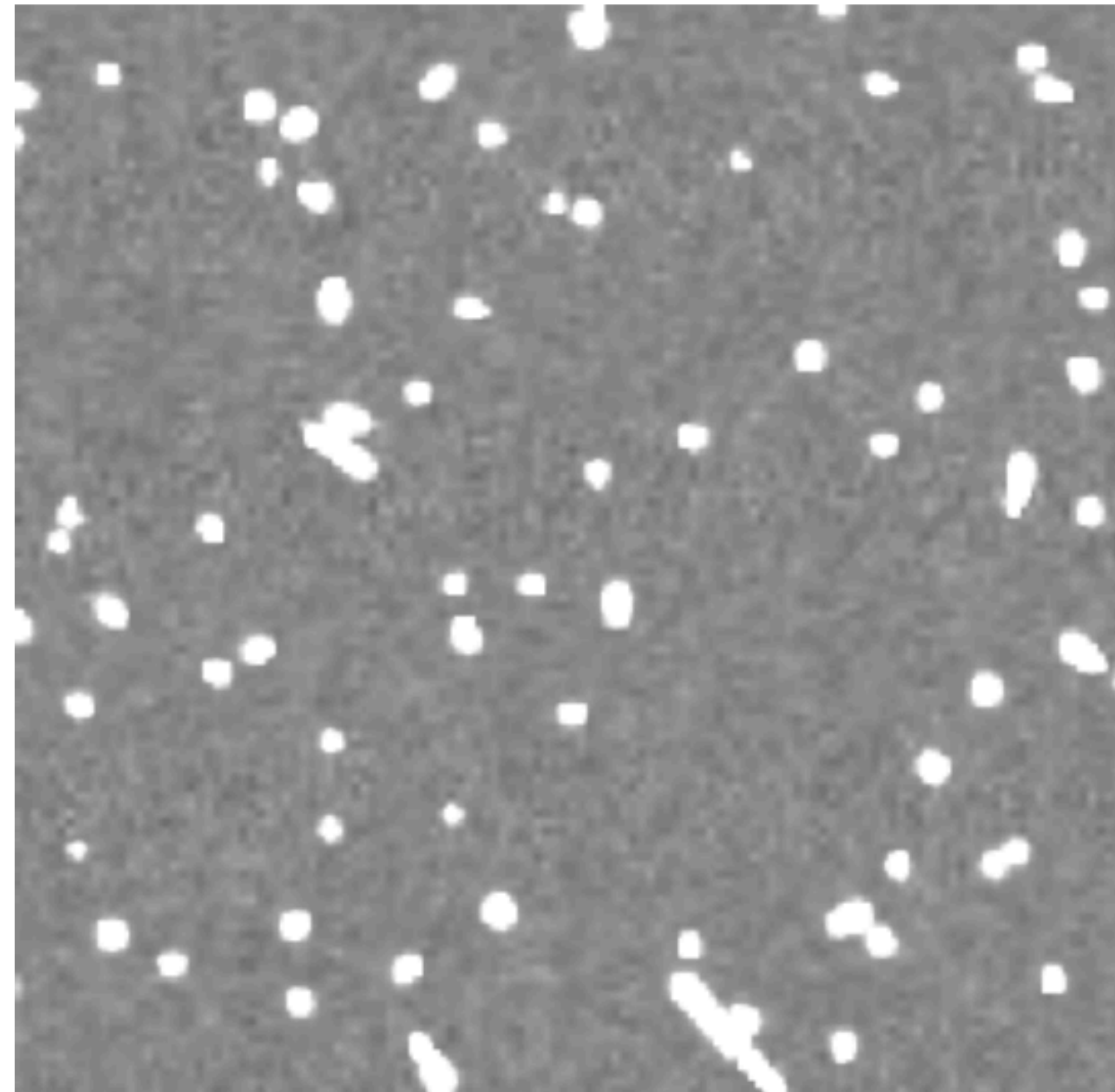
# Noise Generation

- Can sample new noise. Which one is real?

Fake (generated)



Real

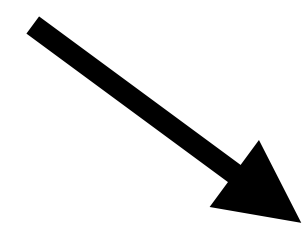


**Besides sampling noise...**

**Inference!**

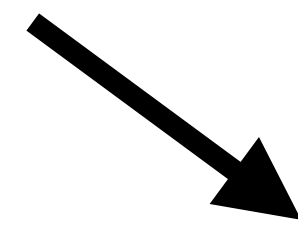


Posterior



$$p(\theta | d) = \frac{p(d | \theta)p(\theta)}{p(d)}$$

Posterior

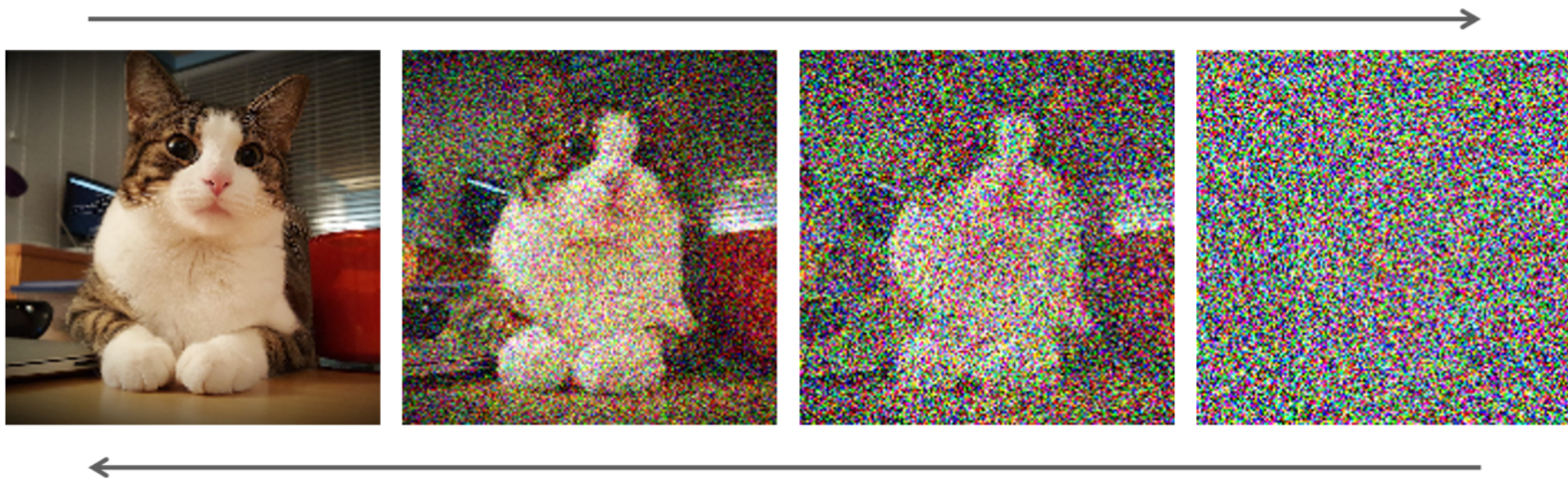


$p(\theta | d)$

Likelihood

$p(d | \theta)p(\theta)$

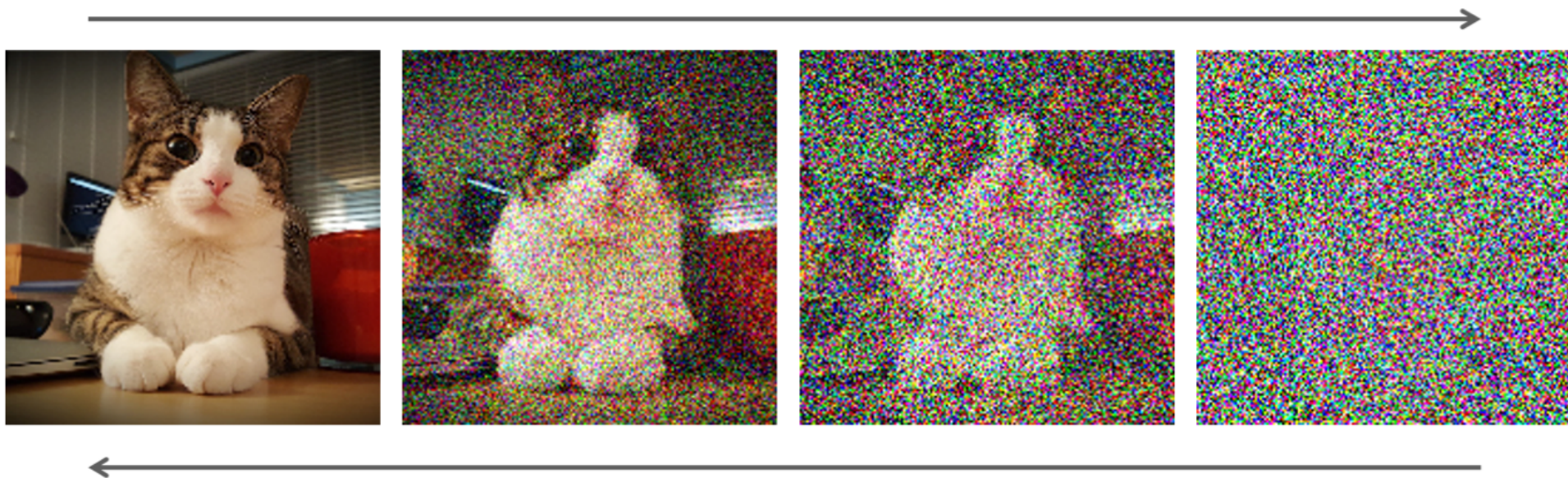
$= \frac{\quad}{p(d)}$



Posterior

Likelihood

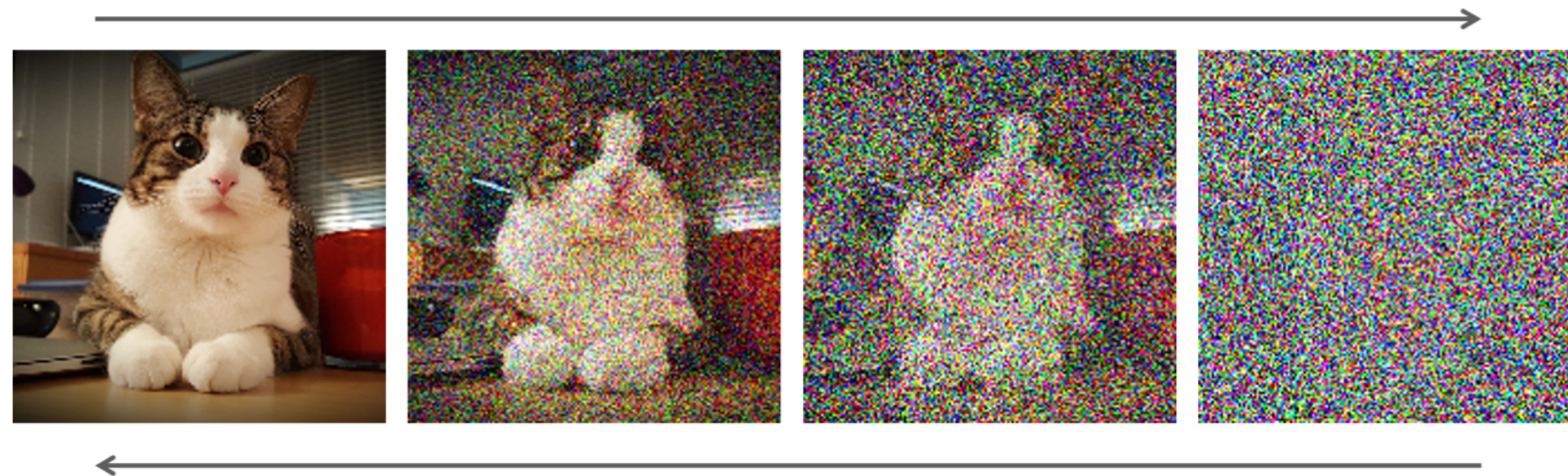
$$\nabla_{\theta} \log p(\theta | d) = \nabla_{\theta} \log p(d | \theta) + \nabla_{\theta} \log p(\theta)$$



Posterior

Likelihood

$$\nabla_{\theta} \log p(\theta | d) = \nabla_{\theta} \log p(d | \theta) + \nabla_{\theta} \log p(\theta)$$



Learned Noise

$$\nabla_x \log Q(x)$$

Posterior

Likelihood

$$\nabla_{\theta} \log p(\theta | d) = \nabla_{\theta} \log p(d | \theta) + \nabla_{\theta} \log p(\theta)$$

# SLIC Framework

## (Score-based Likelihood Characterization)

- Integrate learned noise distribution within well-defined Bayesian framework.
- Currently tested on additive noise  $X = M(\theta) + N$

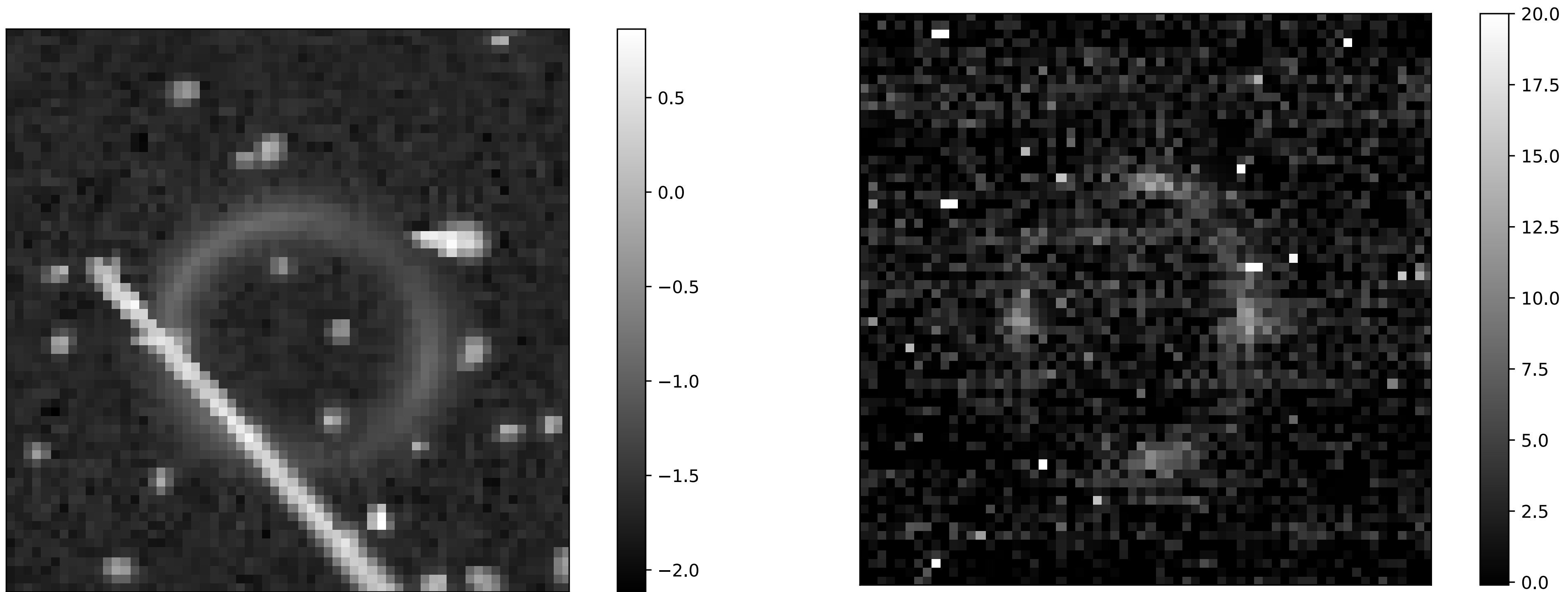
- $P(d | \theta) = Q(d - M(\theta)) \rightarrow \nabla_{\theta} \log Q(d - M(\theta)) = -\nabla \log Q \cdot \nabla_{\theta} M$

Our noise model!

# **SLIC Tests**

# SLIC Example

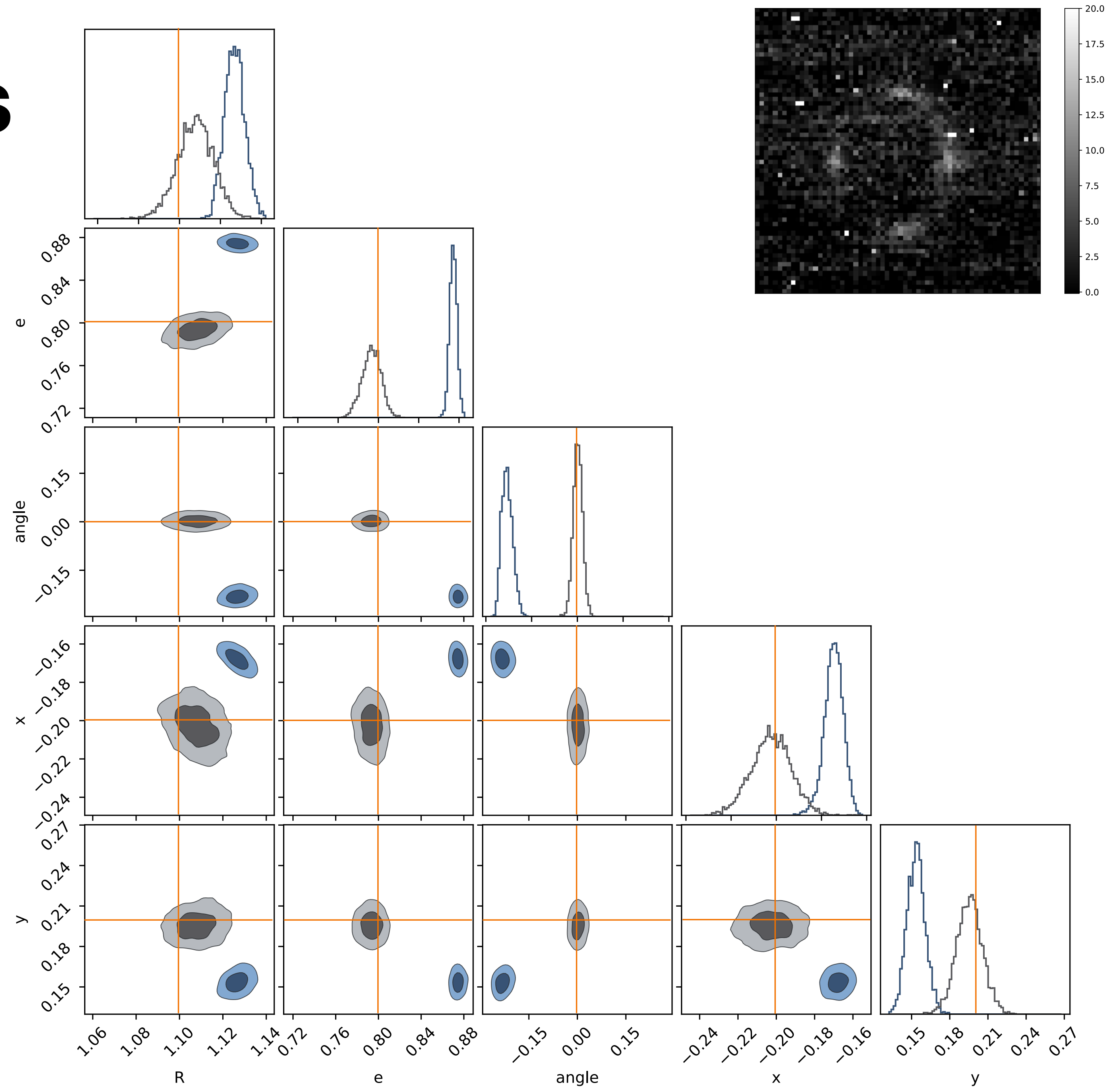
- Real noise + simulated strong lens test problem





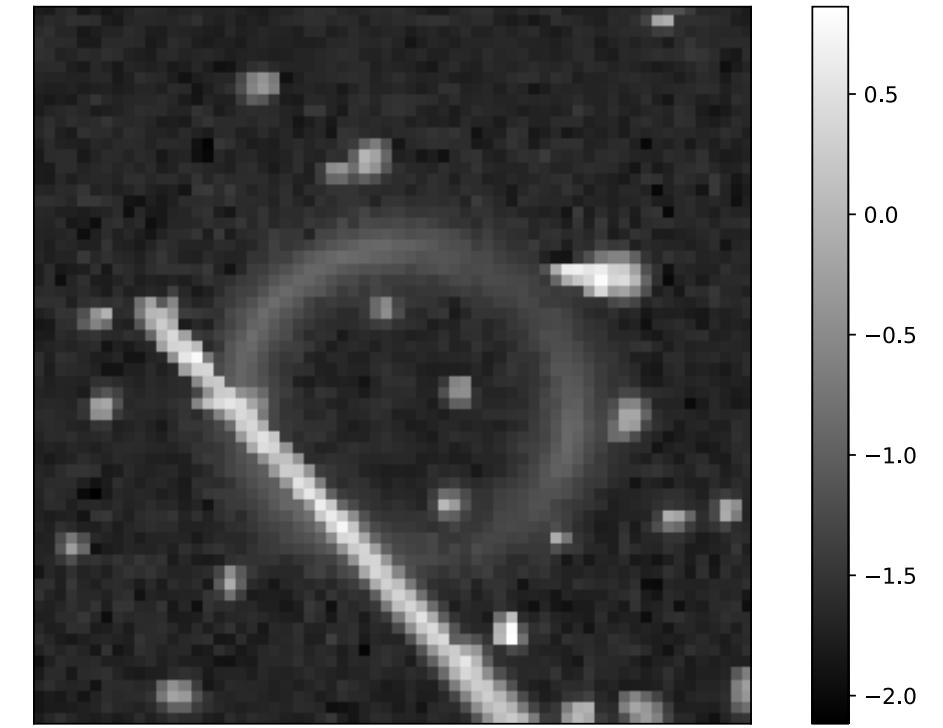
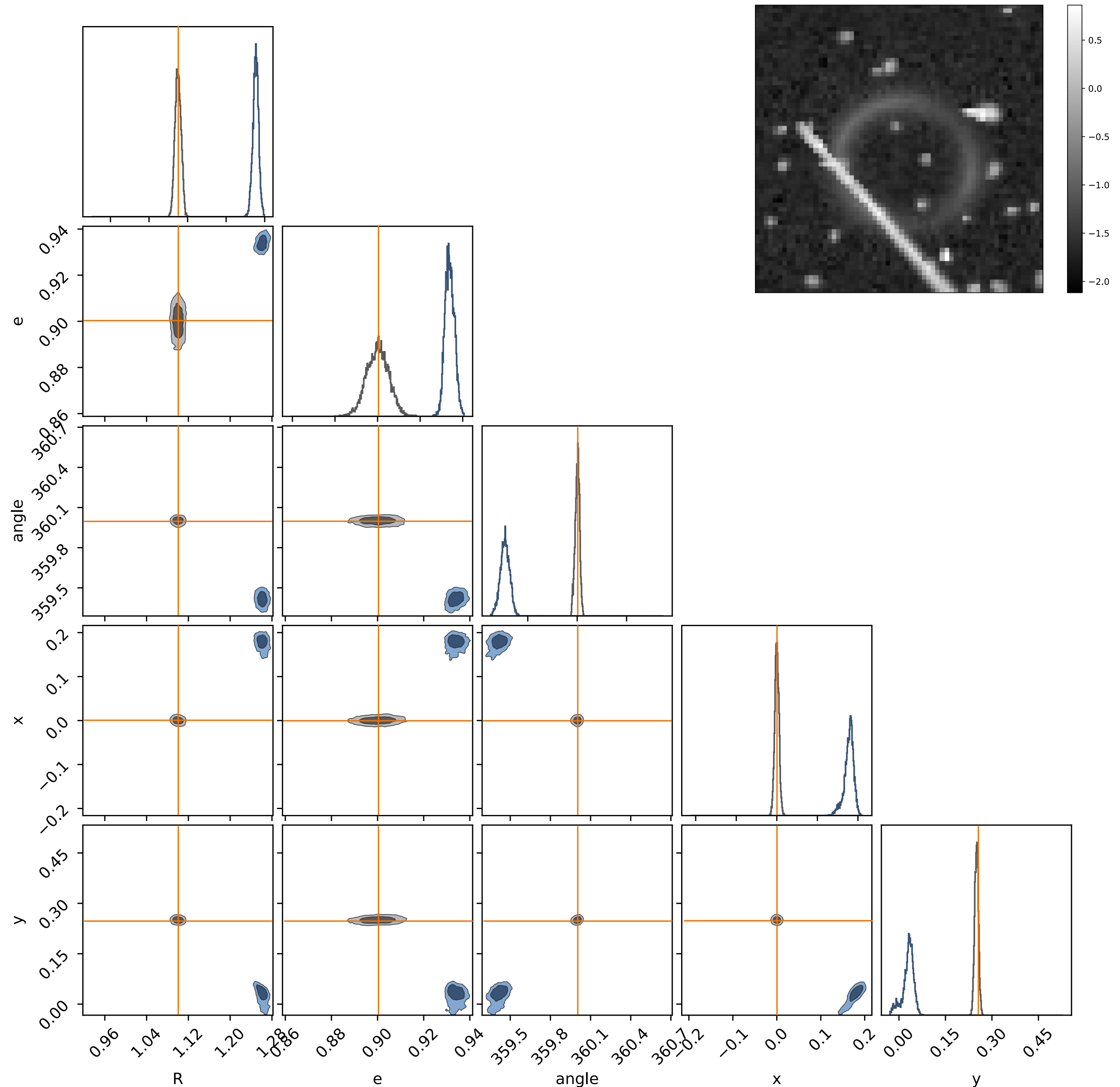
# SLIC JWST Results

- We achieve accurate inference!
- Blue is Gaussian Likelihood
- Grey (middle blob) is SLIC.



# SLIC HST Results

- We achieve accurate inference!
- Blue is Gaussian Likelihood
- Grey (middle blob) is SLIC.



# Beyond Gaussian Noise: A Generalized Approach to Likelihood Analysis with non-Gaussian Noise

RONAN LEGIN,<sup>1,2,3,\*</sup> ALEXANDRE ADAM,<sup>1,2,3,\*</sup> YASHAR HEZAVEH,<sup>1,2,3,4</sup> AND LAURENCE PERREAULT LEVASSEUR<sup>1,2,3,4</sup>

<sup>1</sup>*Department of Physics, Université de Montréal, Montréal, Canada*

<sup>2</sup>*Ciela - Montreal Institute for Astrophysical Data Analysis and Machine Learning, Montréal, Canada*

<sup>3</sup>*Mila - Quebec Artificial Intelligence Institute, Montréal, Canada*

<sup>4</sup>*Center for Computational Astrophysics, Flatiron Institute, 162 5th Avenue, 10010, New York, NY, USA*



# Conclusion

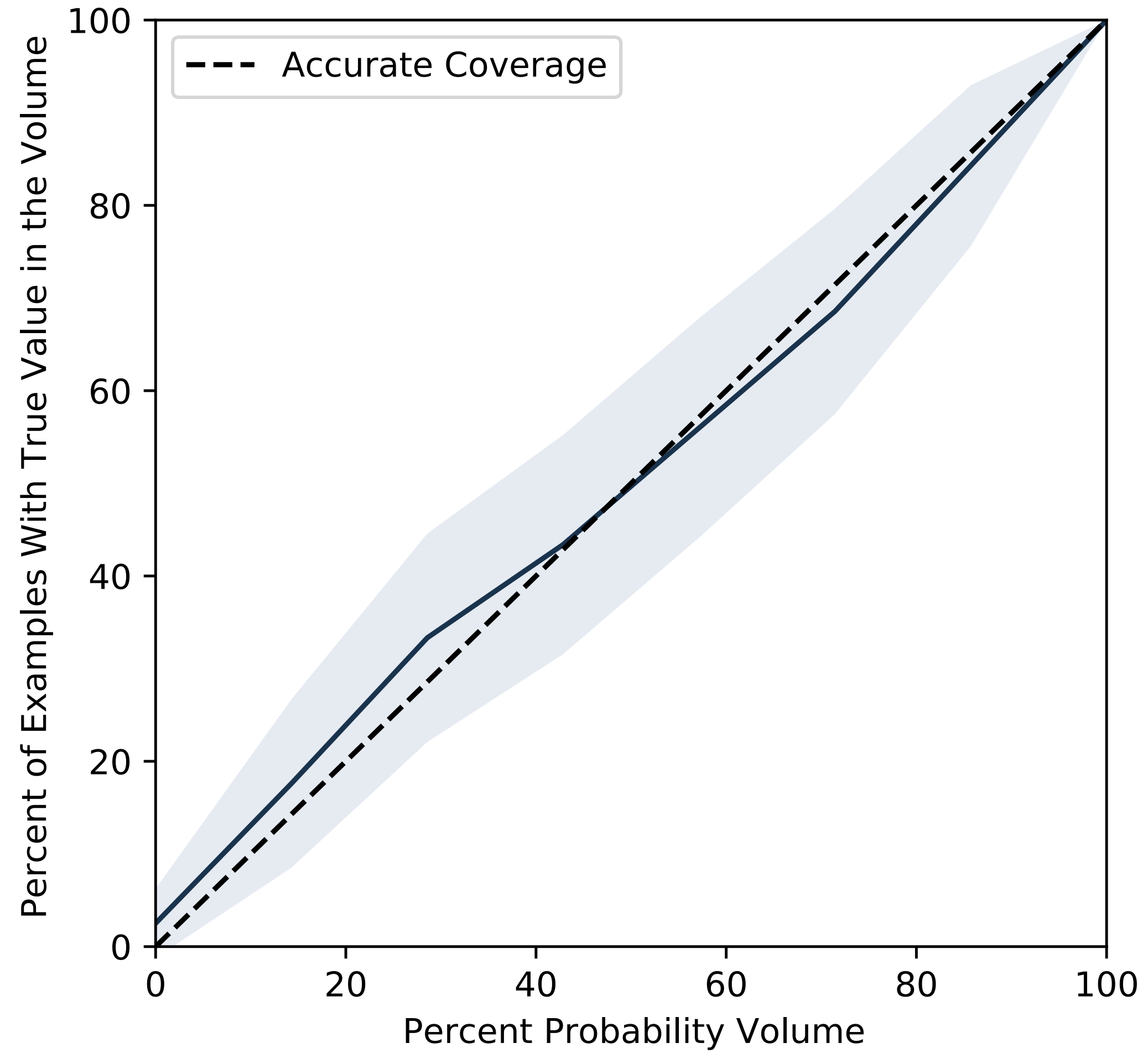
- Astrophysical noise is often non-Gaussian.
- We learn this noise to perform unbiased statistical inference.
- SLIC <https://arxiv.org/abs/2302.03046>

**Thank you!**

**Extra slides**

# Coverage Tests

## SLIC



# The Machine Learning

- Score matching with transfer kernel  $p(x_t | x_0) = \mathcal{N}(x_0, \sigma^2(t))$

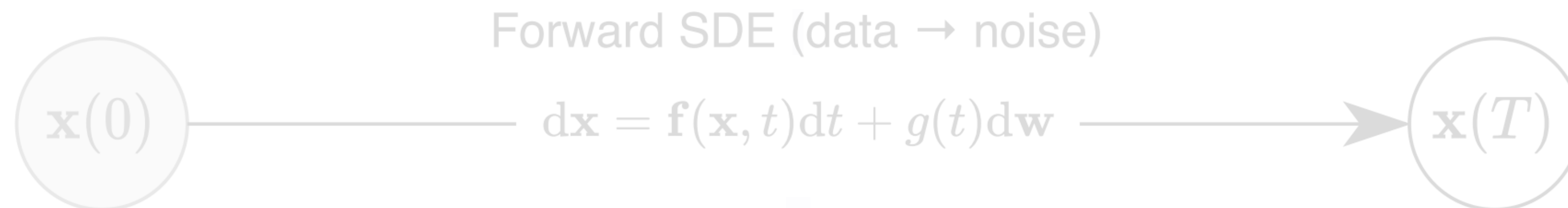
$$\|f_\eta(x_t, t) - \nabla_{x_t} \log p(x_t | x_0)\|^2$$

# Posterior sampling

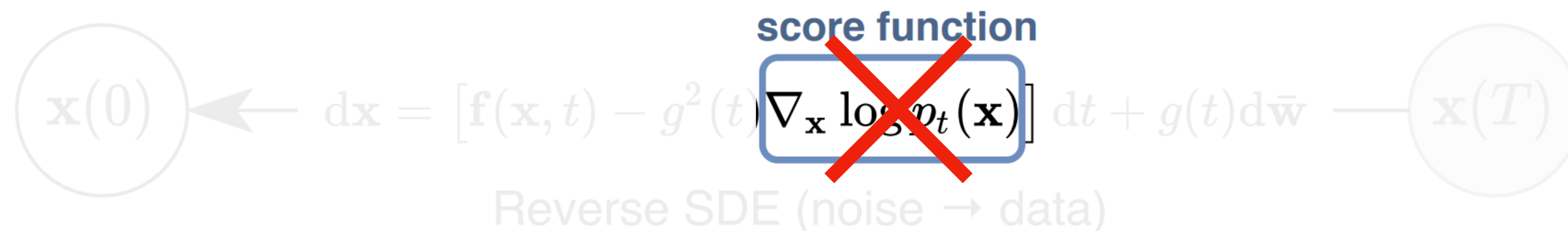
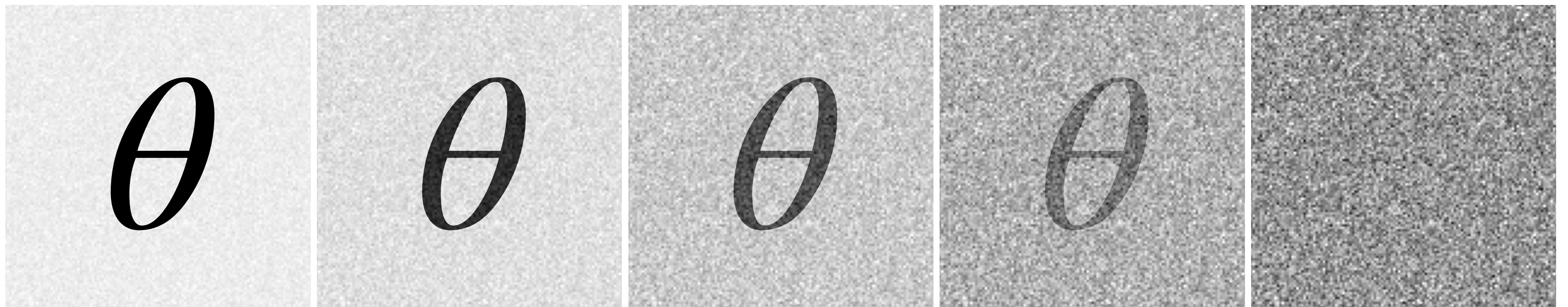
Likelihood

Prior

$$\nabla_{\theta} \log P_t(X | \theta) + \nabla_{\theta} \log P_t(\theta)$$



$\theta$



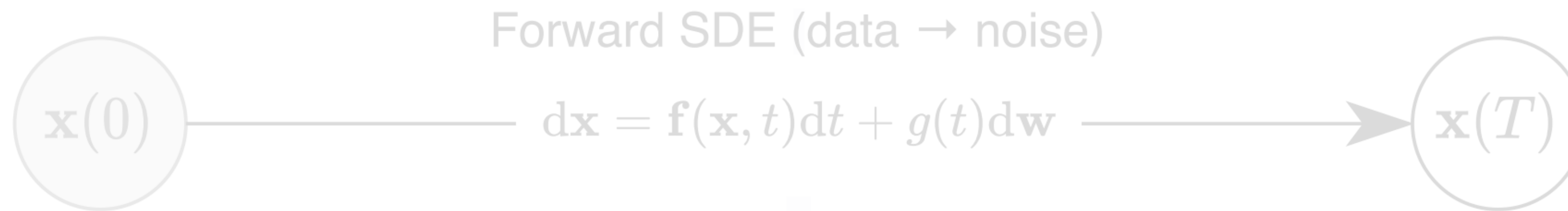


# SLIC inference

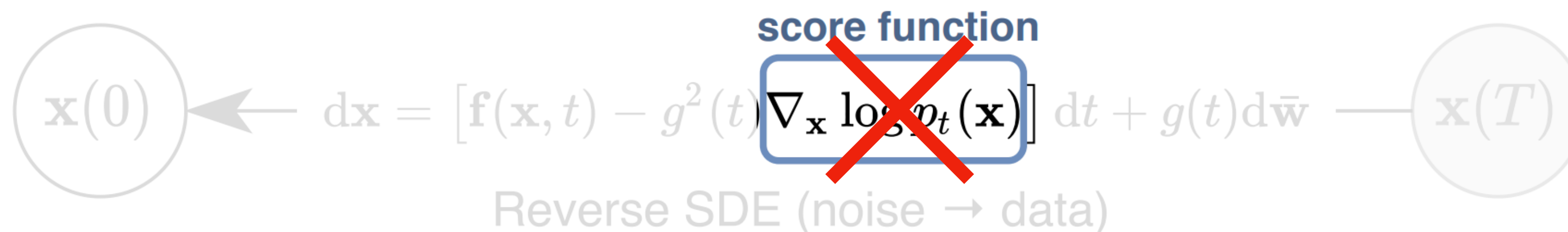
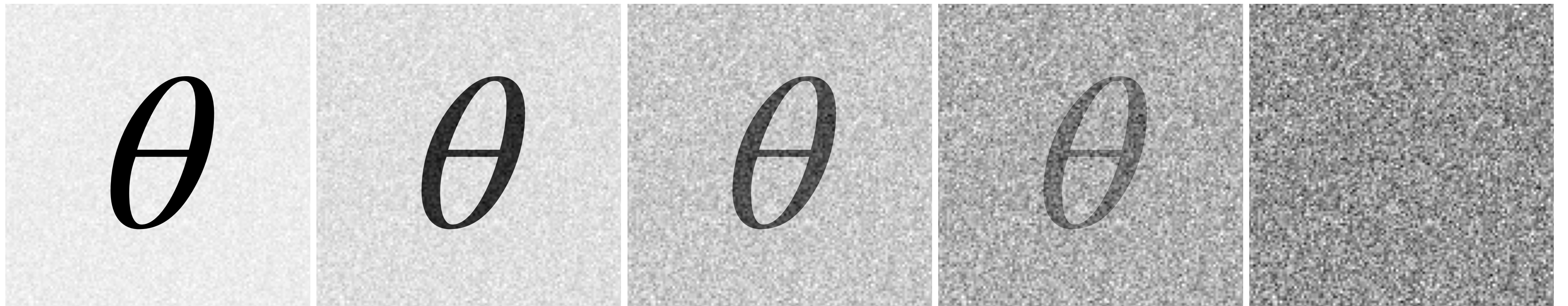
Likelihood

Prior

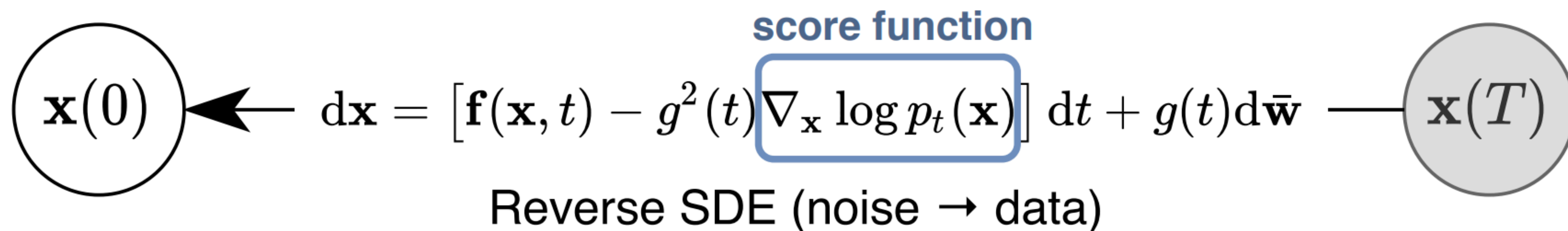
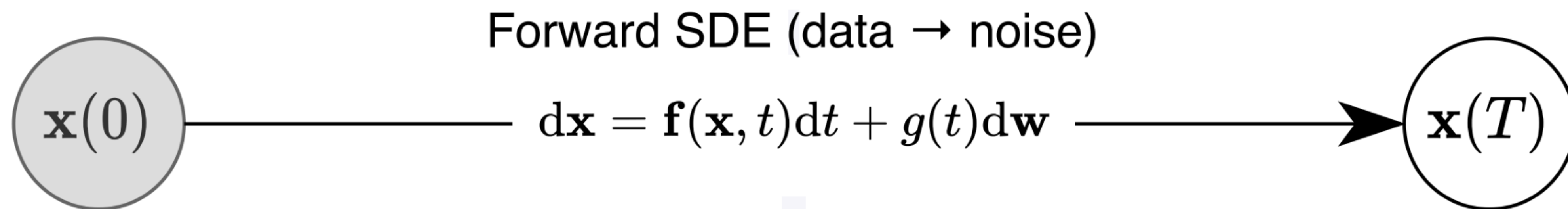
$$-\nabla_{\mathbf{x}-M(\theta)} \log Q \cdot \nabla_{\theta} M(\theta) + \nabla_{\theta} \log P_t(\theta)$$



$\theta$



# Score-based generative models



# SLIC Framework

## (Score-based Likelihood Characterization)

- Assume additive noise  $X = M(\theta) + N$
- Given previous point, we can write likelihood  $P(X | \theta) = Q(X - M(\theta))$
- $Q(X - M(\theta))$  is probability density of noise

# SLIC trick

## Decomposition with chain rule

- $\nabla_{\theta} \log Q(X - M(\theta)) = - \nabla_{X - M(\theta)} \log Q(X - M(\theta)) \cdot \nabla_{\theta} M(\theta)$
- This separates noise distribution  $Q$  from forward simulator  $M$

# SLIC trick

## Decomposition with chain rule

- $\nabla_{\theta} \log Q(X - M(\theta)) = - \nabla_{X - M(\theta)} \log Q(X - M(\theta)) \cdot \nabla_{\theta} M(\theta)$
- This separates noise distribution  $Q$  from forward simulator  $M$

Model  $\nabla_{X - M(\theta)} \log Q(X - M(\theta))$  using score network!