

Space weather impacts and predictions: relevant spatial and temporal scales

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 - Power transmission industry.
 - Human spaceflight.
- Addressing the end-user needs (and dealing with the spatiotemporal complexity of the relevant space environmental phenomena).
 - Geomagnetic storm benchmarks.
 - Solar energetic particle (SEP) predictions.



Identification of end-user needs



End-user needs

- If we acknowledge space weather as the societal dimension of heliophysics, understanding the impacts and associated end-user needs are the foundation of the field.
- Our space weather work must be informed by those needs and strive toward generating information that is actionable.





End-user needs – power transmission industry

- While also predictions are of interest, the main U.S. focus right now is on hazard assessments.
- To enable hazard assessments, space weather extremes need to be communicated to the end-user in the form of benchmarks.





GMD benchmark requirements

- Science side needs to provide information about a physical parameter that is directly applicable/actionable to further engineering analyses. (geoelectric field)
- We need to address the following key characteristics of the extreme geoelectric fields:
 - (Element 1) Amplitude. i. .
 - (Element 2) Spatial structure. ii.
 - iii. Temporal waveform.
 - Full 1-3 day storm characterized.
 - 1-10 s sampling to capture rapid enhancements that may compromise voltage stability.
 - Longer duration enhancements necessary for thermal heatingrelated problems.

(Element 3)

 Science analyses also need to characterize the occurrence rates of i-iii.



Pulkkinen et al. (2012)





GMD benchmark requirements

- The geomagnetic induction process that generates the geoelectric field is dependent on external and internal factors:
 - iv. Many different near space electric currents systems contribute to driving of geomagnetic induction. The effect of the geomagnetic latitude, and possibly local time, needs to be taken into account. (Element 4)
 - v. The local ground conductivity dictates the ground response. Local geology needs to be taken into account. (Element 5)



End-user needs – human spaceflight

- While low-inclination LEO (ISS orbit) is fairly benign from the space radiation perspective, deep space environment experienced in the Artemis program poses a much more significant challenge.
- The key problem is ionizing radiation: > 10 MeV ions for EVAs and > 100 MeV ions for the crew inside the vehicle.
- Primary sources for energetic ions contributing to possible problems include galactic cosmic rays, SEPs and inner radiation belt – only the SEP component discussed here.





End-user needs – human spaceflight

- Due to the SEP challenge, Artemis will have storm shelter as a part of the ops. The shelter needs to be deployed in 30 min from (*Townsend et al.*, 2018) → Predictive capability plays a critical role in the ops.
- We need to have information about elevated, likely mostly CME shock-driven, energetic ion fluxes at the location of the vehicle.





End-user needs – human spaceflight





Addressing the end-user needs



GMD benchmark(s) – spatiotemporal representation per the NERC standard

$$\overline{E}(x, y, t) = \begin{bmatrix} E_x(x, y, t) \\ E_y(x, y, t) \end{bmatrix}$$

 $\approx E_{peak}(x,y) \begin{bmatrix} f_x(t)g_x(x,y) \\ f_v(t)g_v(x,y) \end{bmatrix}$

 $\overline{E}(x, y, t)$ depends on:

- External excitation $\bar{B}_{ext}(x, y, t)$
- Ground response dictated by $\sigma(x, y, z)$

Assume spatially homogeneous field

 $\approx E_{peak}(x,y) \begin{bmatrix} f_x(t) \\ f_y(t) \end{bmatrix} \cdot 1$ Factorize & approximate the

primary dependencies

$$\approx E_0 \cdot \alpha(y) \cdot \beta(x, y) \begin{bmatrix} f_x(t) \\ f_y(t) \end{bmatrix} \cdot 1$$

Latitude dependence

Ground response dependence



GMD benchmark(s) – regional vs localized enhancements

Pulkkinen et al. (2015)





 E_0 quantified with a spatial average E-field applied regionally

E_0 quantified with individual stations

E-field applied locally



GMD benchmark(s)

- Element 1: amplitude E_0
- Element 2: spatial structure
- Element 3: reference temporal waveform f(t)
- Element 4: geomagnetic latitude dependence $\alpha(y)$
- Element 5: dependence on the local ground conductivity $\beta(x,y)$





100 years

values

of 10 s 1

Return Leve



00:00 03:00 06:00 09:00 12:00 15:00 18:00 21:00 00:00 03:00 06:00 Time (UT) NERC GMD benchmark white paper



SEP prediction approaches

Models available at iswa.gsfc.nasa.gov & ccmc.gsfc.nasa.gov



SOHO/Costep Proton Flux Forecast

RELEASE proton flux forecast at CCMC (data source: costep2) by ETPH IEAP CAU Kiel and SWRI – data gaps due to limited DSN coverage

9-15.8 MeV

4-9 MeV

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15.8-39.8 MeV 28.2-50.1 MeV



Conclusions

- From the space weather standpoint, end-user impacts and needs are the fundamental driver for identifying i) actionable physical parameters of interest, ii) relevant spatiotemporal scales.
- "Unfortunately" it is often necessary to address a blend of global and local spatial scales and a wide range of temporal scales – space weather challenges our understanding of the heliophysics system.
- Empirical, first-principles, handwaving etc. approaches all being used – the nature of the approach does not matter as long as it works.
- It is not all about predictions: In some applications general characterization of extreme environments is currently of greater interest.



Backup

Integrated Exploration Manifest: 2019-2024



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